

OVERVIEW: UDIG Workshop BNL

Stephen Parke
Fermilab
Oct 16, 2008

LBL Neutrino Oscillations
Proton Decay
Supernova Neutrinos
Other: Solar & Atm. Nus ...

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"All the News
That's Fit to Print"

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Mass Found in Elusive Particle; Universe May Never Be the Same

**Discovery on Neutrino
Rattles Basic Theory
About All Matter**

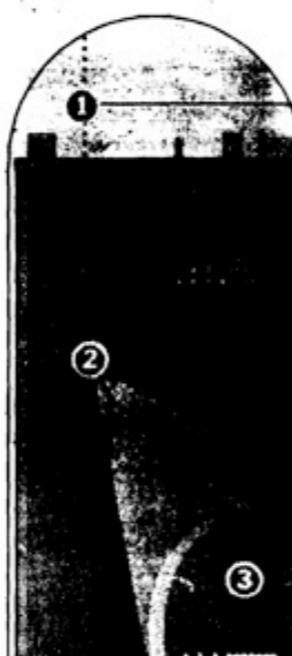
By MALCOLM W. BROWNE

TAKAYAMA, Japan, June 5 — In what colleagues hailed as a historic landmark, 120 physicists from 23 research institutions in Japan and the United States announced today that they had found the existence of mass in a notoriously elusive subatomic particle called the neutrino.

The neutrino, a particle that carries no electric charge, is so light that it was assumed for many years to have no mass at all. After today's announcement, cosmologists will have to confront the possibility that much of the mass of the universe is in the form of neutrinos. The discovery will also compel scientists to revise a highly successful theory of the composition of matter known as the Standard Model.

Word of the discovery had drawn some 300 physicists here to discuss neutrino research. Among other things, they said, the finding of neutrino mass might affect theories about the formation and evolution of galaxies and the ultimate fate of the

Detecting Neutrinos



Neutrinos pass through the Earth's surface to a tank filled with 12.5 million gallons of ultra-pure water . . .

. . . and collide with other particles . . .

. . . producing a cone-shaped flash of light.

The light is recorded by 11,200 20-inch light amplifiers that cover the inside of the tank.

And Detecting Their Mass

By analyzing the cones of light,

OKLAHOMA BLAST BRINGS LIFE TERM FOR TERRY NICHOLS

'ENEMY OF CONSTITUTION'

Judge Denounces Conspiracy
and Hears From the Victims
of a Terrifying Ordeal

By JO THOMAS

DENVER, June 4 — Calling him "an enemy of the Constitution," a Federal judge today sentenced Terry L. Nichols to life in prison without the possibility of parole for conspiring to bomb the Oklahoma City Federal Building, the deadliest terrorist attack ever on American soil.

In passing sentence after hearing from survivors of the blast and relatives of some of the 168 people who died in it, the judge, Richard P. Matsch of Federal District Court, said, "This was not a murder case."

He added: "It is a crime and the victims have spoken eloquently here. But it is not a crime as to them so much as it is a crime against the Constitution of the United States. That's the victim."



Bajram Curri, in nc

Mixing Matrix:

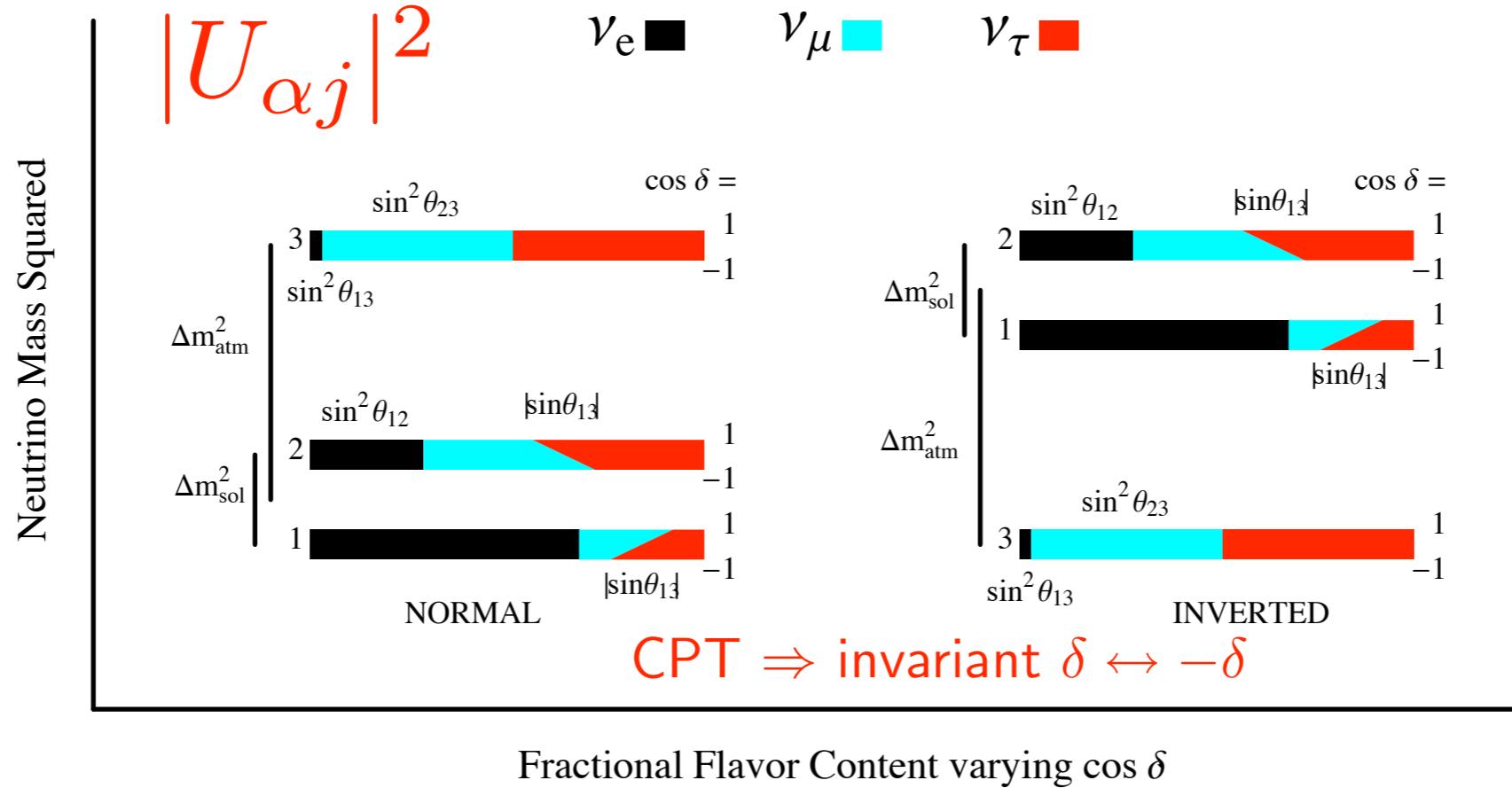
$$|\nu_e, \nu_\mu, \nu_\tau\rangle_{flavor}^T = U_{\alpha i} |\nu_1, \nu_2, \nu_3\rangle_{mass}^T$$

$$U_{\alpha i} = \begin{pmatrix} 1 & & \\ c_{23} & s_{23} & \\ -s_{23} & c_{23} & \end{pmatrix} \begin{pmatrix} c_{13} & & s_{13}e^{-i\delta} & \\ & 1 & & \\ & -s_{13}e^{i\delta} & & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & & s_{12} & \\ -s_{12} & c_{12} & & \\ & & 1 & \end{pmatrix} \begin{pmatrix} 1 & & e^{i\alpha} & \\ & e^{i\beta} & & \end{pmatrix}$$

Atmos. L/E $\mu \rightarrow \tau$ Atmos. L/E $\mu \leftrightarrow e$ Solar L/E $e \rightarrow \mu, \tau$ $0\nu\beta\beta$ decay

500km/GeV

15km/MeV



$$\delta m_{sol}^2 = +7.6 \times 10^{-5} \text{ eV}^2$$

$$|\delta m_{atm}^2| = 2.4 \times 10^{-3} \text{ eV}^2$$

$$|\delta m_{sol}^2| / |\delta m_{atm}^2| \approx 0.03$$

$$\sqrt{\delta m_{atm}^2} = 0.05 \text{ eV} < \sum m_{\nu_i} < 0.5 \text{ eV} = 10^{-6} * m_e$$

$$\sin^2 \theta_{12} \sim 1/3$$

$$\sin^2 \theta_{23} \sim 1/2$$

$$\sin^2 \theta_{13} < 3\%$$

$$0 \leq \delta < 2\pi$$

One Global Fit:

Dominated by

parameter	best fit	2σ	3σ
Δm_{21}^2 [10^{-5} eV 2]	$7.65^{+0.23}_{-0.20}$	7.25–8.11	7.05–8.34
$ \Delta m_{31}^2 $ [10^{-3} eV 2]	$2.40^{+0.12}_{-0.11}$	2.18–2.64	2.07–2.75
$\sin^2 \theta_{12}$	$0.304^{+0.022}_{-0.016}$	0.27–0.35	0.25–0.37
$\sin^2 \theta_{23}$	$0.50^{+0.07}_{-0.06}$	0.39–0.63	0.36–0.67
$\sin^2 \theta_{13}$	$0.01^{+0.016}_{-0.011}$	≤ 0.040	≤ 0.056

KamLAND
MINOS
SNO
SuperK
Chooz

arXiv:0808.2016

Neutrino

Mass Spectrum:

- Quasi-Degenerate ?
- Hierarchical ?
- Normal or Inverted ?

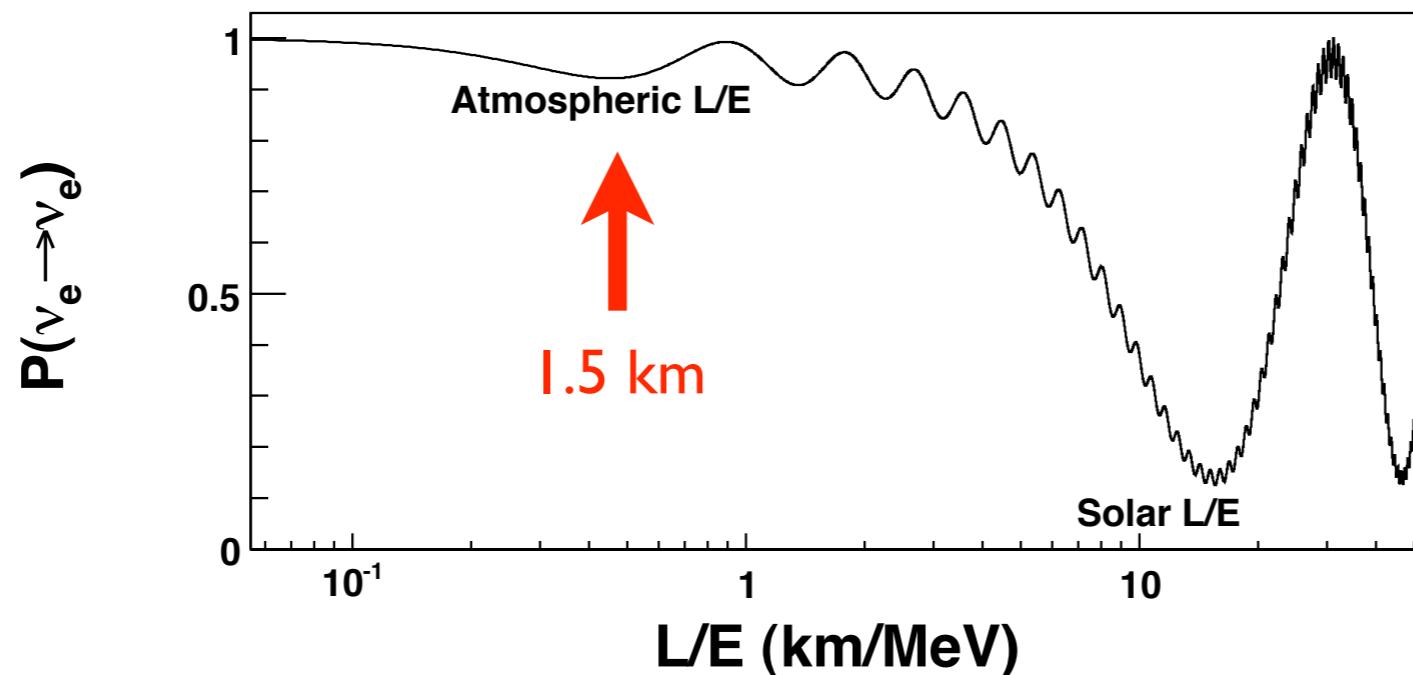
Mixings:

- Deviations from $U_{Tri-Bi-Max}$
 $\sin^2 \theta_{13}$, $(\sin^2 \theta_{23} - 1/2)$, $(\sin^2 \theta_{12} - 1/3)$
- Relationship between these deviations and
 $V_{CKM} - 1$
if any ?
- Magnitude and sign of CPV:
 $\propto \sin \theta_{13} \sin \delta$

$\sin^2 \theta_{13}$ from Reactor Neutrinos:

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta_{21} - \sin^2 2\theta_{13} (\cos^2 \theta_{12} \sin^2 \Delta_{31} + \sin^2 \theta_{12} \sin^2 \Delta_{32})$$

kinematic phase:
 $\Delta_{ij} \equiv \frac{\delta m_{ij}^2 L}{4E}$

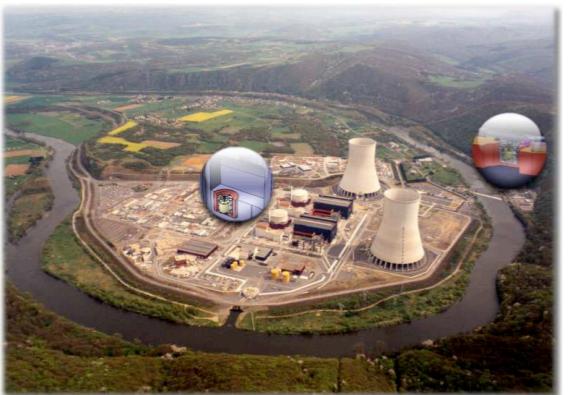


$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\delta m_{ee}^2 L}{4E} \right) - \mathcal{O}(\Delta_{21})^2$$

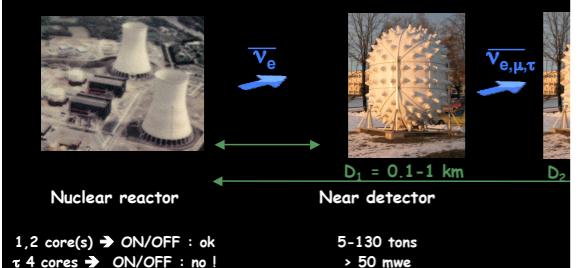
< 0.002

> 0.01 $\delta m_{ee}^2 = \cos^2 \theta_{12} |\delta m_{31}^2| + \sin^2 \theta_{12} |\delta m_{32}^2|$

Double
Chooz:



One nuclear plant & two det

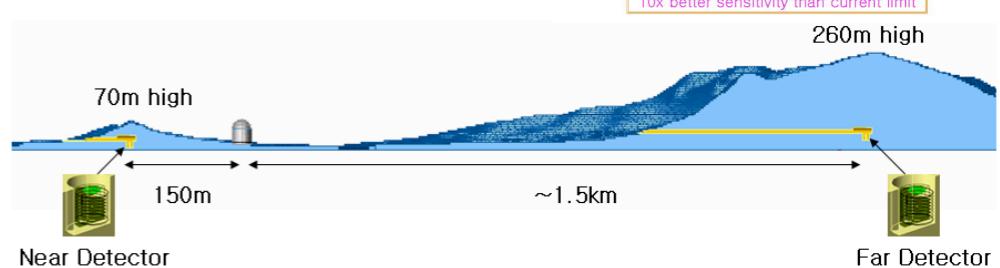
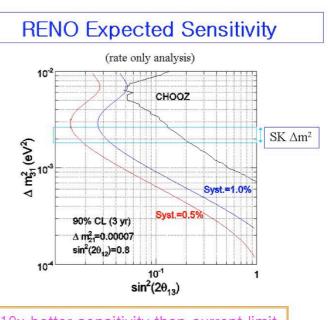
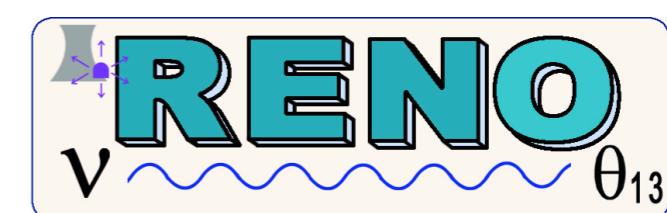


Daya Bay



push the limi
 $\sin^2 2\theta_{13} < 0$

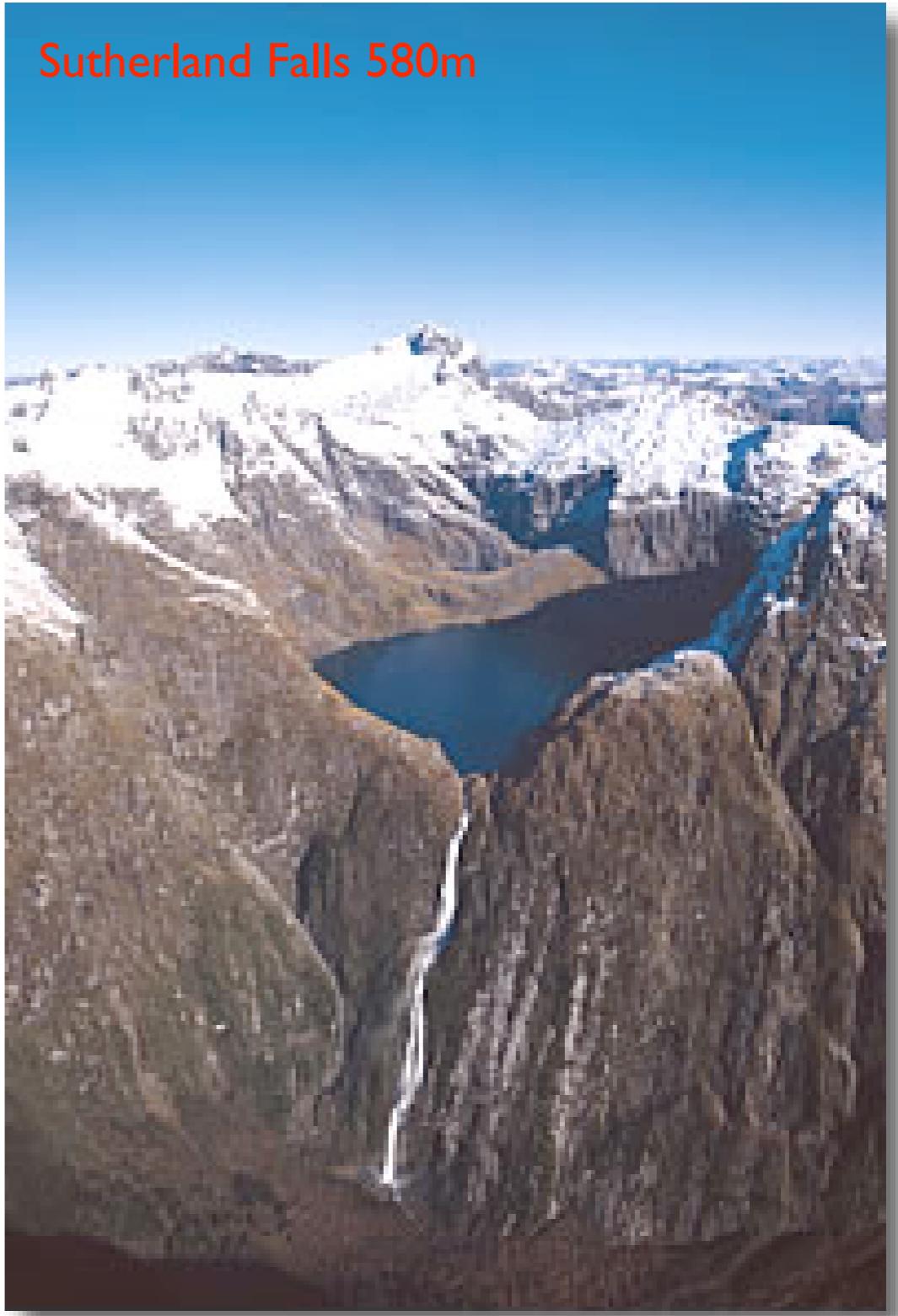
(Reactor Experiment for Neutrino Oscillation)

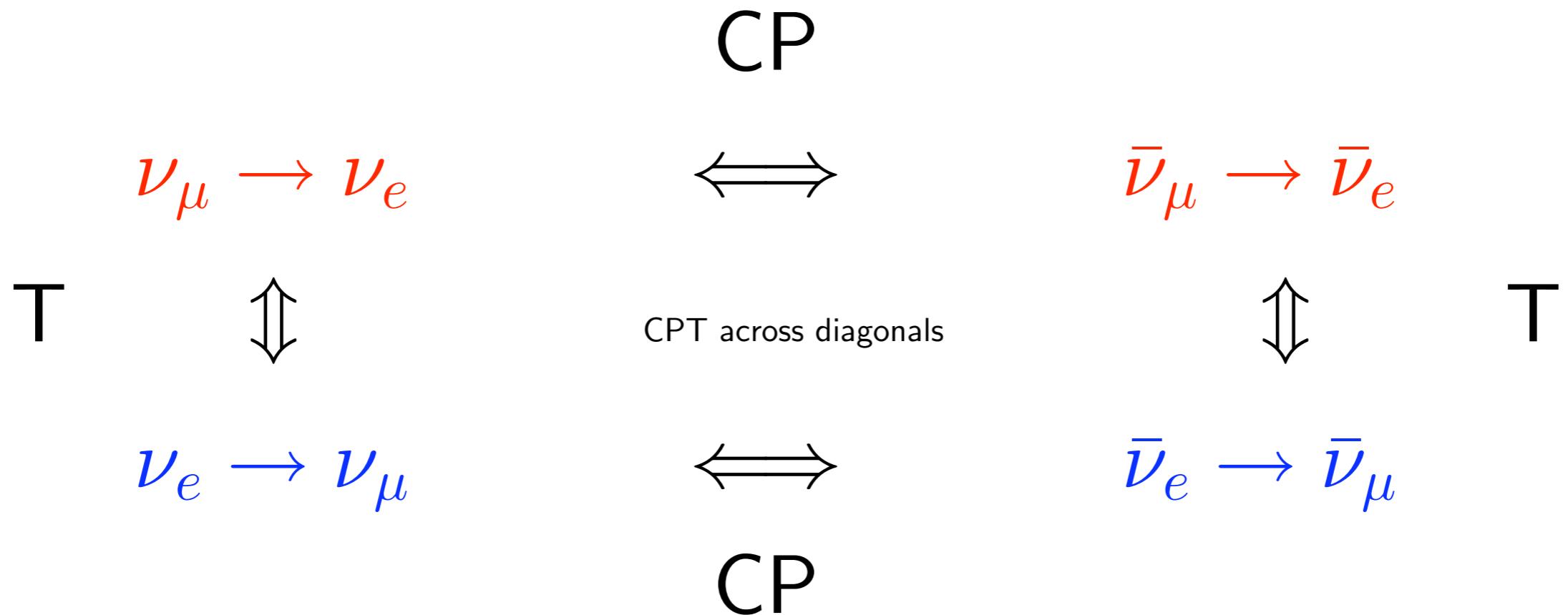


$\sin^2 \theta_{13}$ from LBL:

$$\nu_\mu \rightarrow \nu_e$$

and related processes:



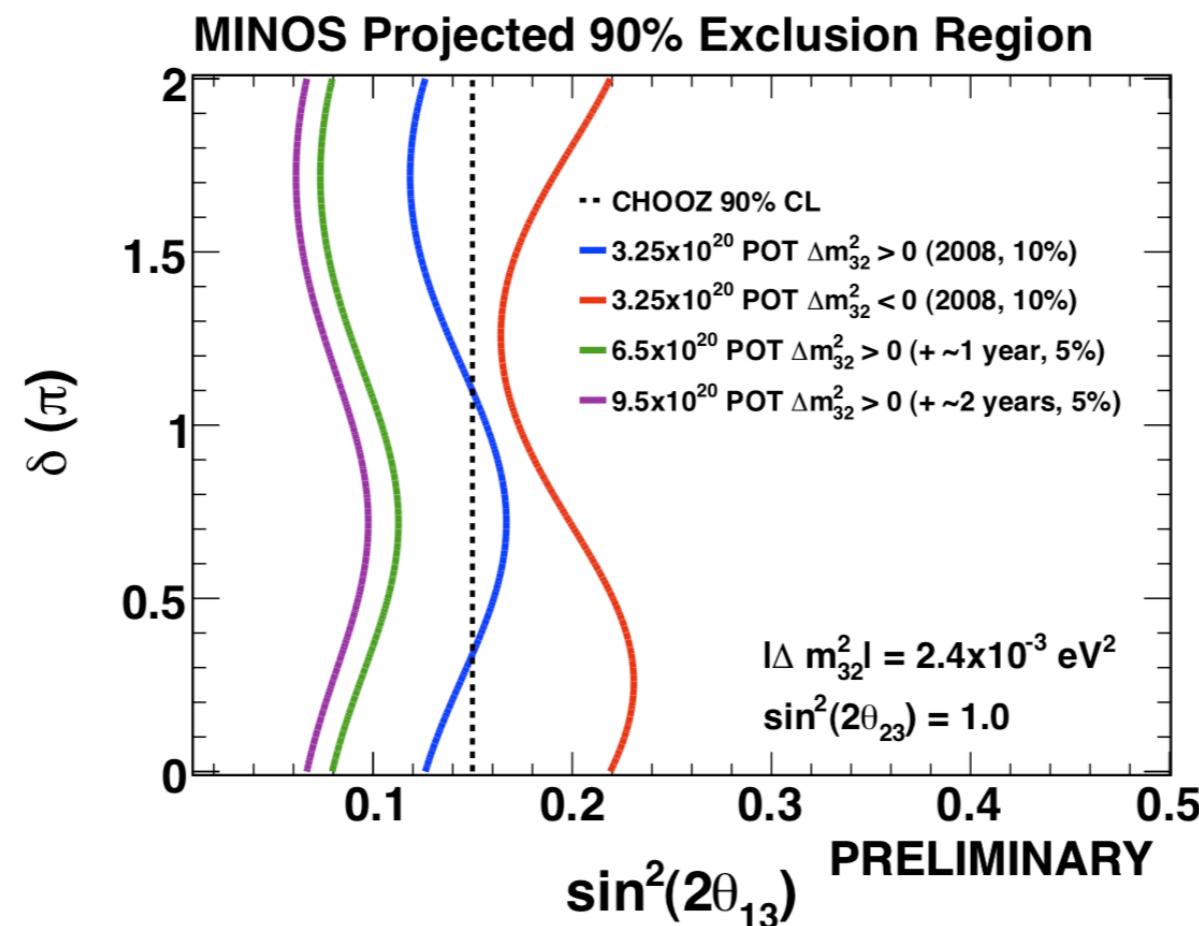


- First Row: Superbeams where ν_e contamination $\sim 1\%$
- Second Row: ν -Factory or β -Beams, no beam contamination

MINOS:

ν_e Sensitivity

H. Gallagher
Tufts University
Neutrino 2008
May 27, 2008



At CHOOZ limit expect 12 ν_e signal events and 42 background events with 3.25×10^{20} protons.

Use sidebands to study predicted far detector backgrounds.

Vacuum LBL:

$$\nu_\mu \rightarrow \nu_e$$

$$P_{\mu \rightarrow e} \approx | \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} |^2$$

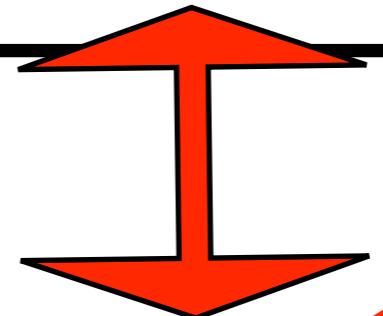
$$\Delta_{ij} = \delta m_{ij}^2 L / 4E$$

CP violation !!!

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \sin \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \sin \Delta_{21}$

$$P_{\mu \rightarrow e} \approx P_{atm} + 2\sqrt{P_{atm}P_{sol}} \cos(\Delta_{32} \pm \delta) + P_{sol}$$

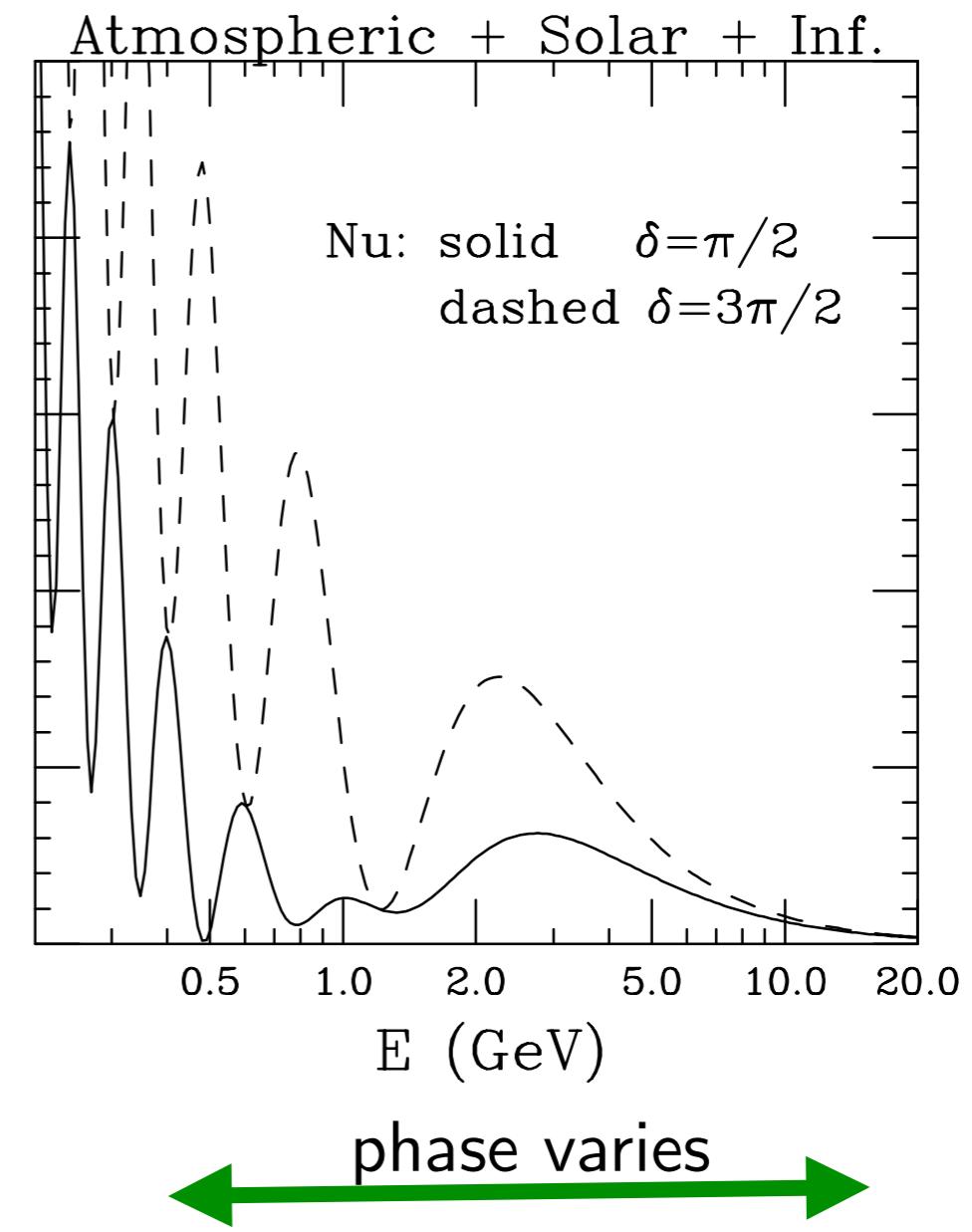
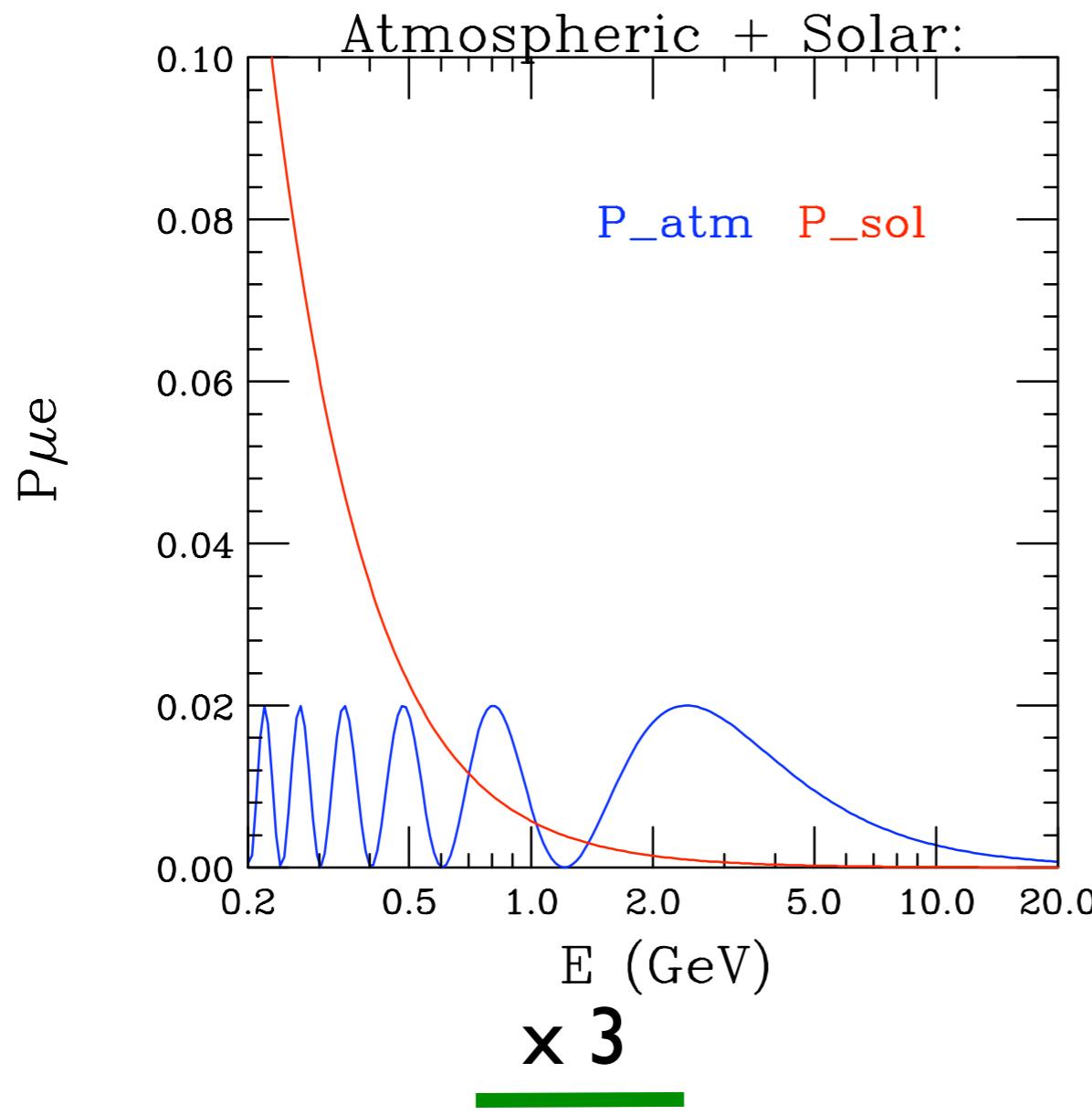


only CPV

$$\cos(\Delta_{32} \pm \delta) = \cos \Delta_{32} \cos \delta \mp \sin \Delta_{32} \sin \delta$$

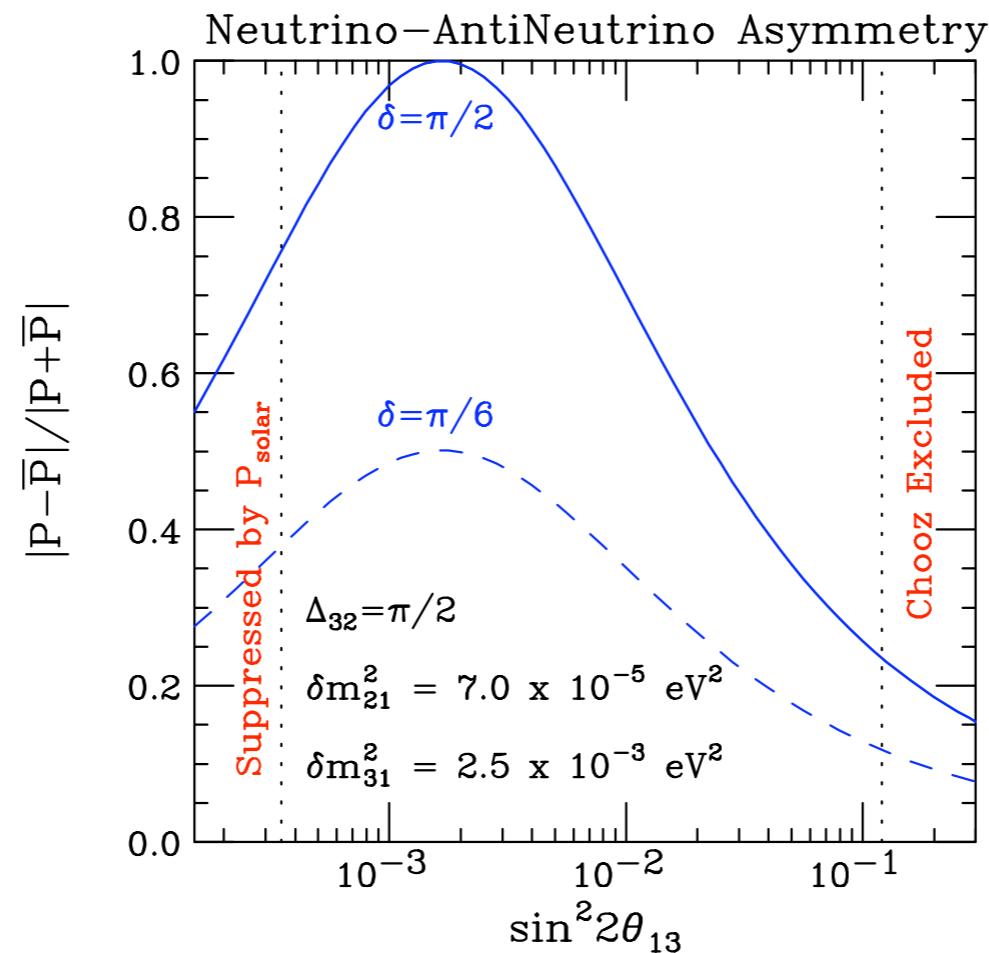
$$P(\nu_\mu \rightarrow \nu_e) \approx |\sqrt{P_{atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{sol}}|^2$$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$



$$P_{\mu \rightarrow e} \approx | \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} |^2$$

Asymmetry Peaks:



$$P_{atm} \leq P_{sol} \quad \text{when} \quad \sin^2 2\theta_{13} \leq \frac{\sin^2 2\theta_{12}}{\tan^2 \theta_{23}} \left(\frac{\delta m_{21}^2}{\delta m_{31}^2} \right)^2 \approx 0.001$$

In Matter:

$$P_{\mu \rightarrow e} \approx | \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} |^2$$

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

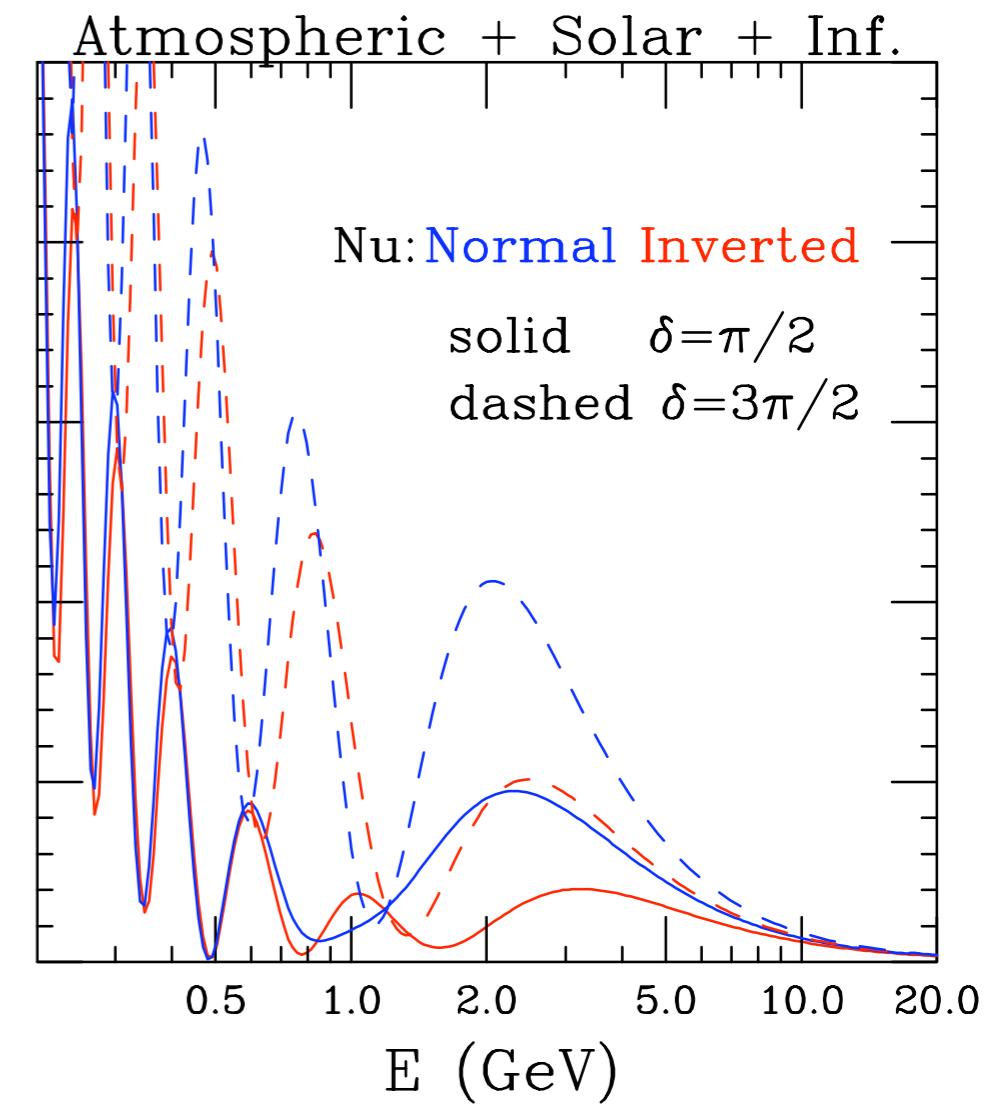
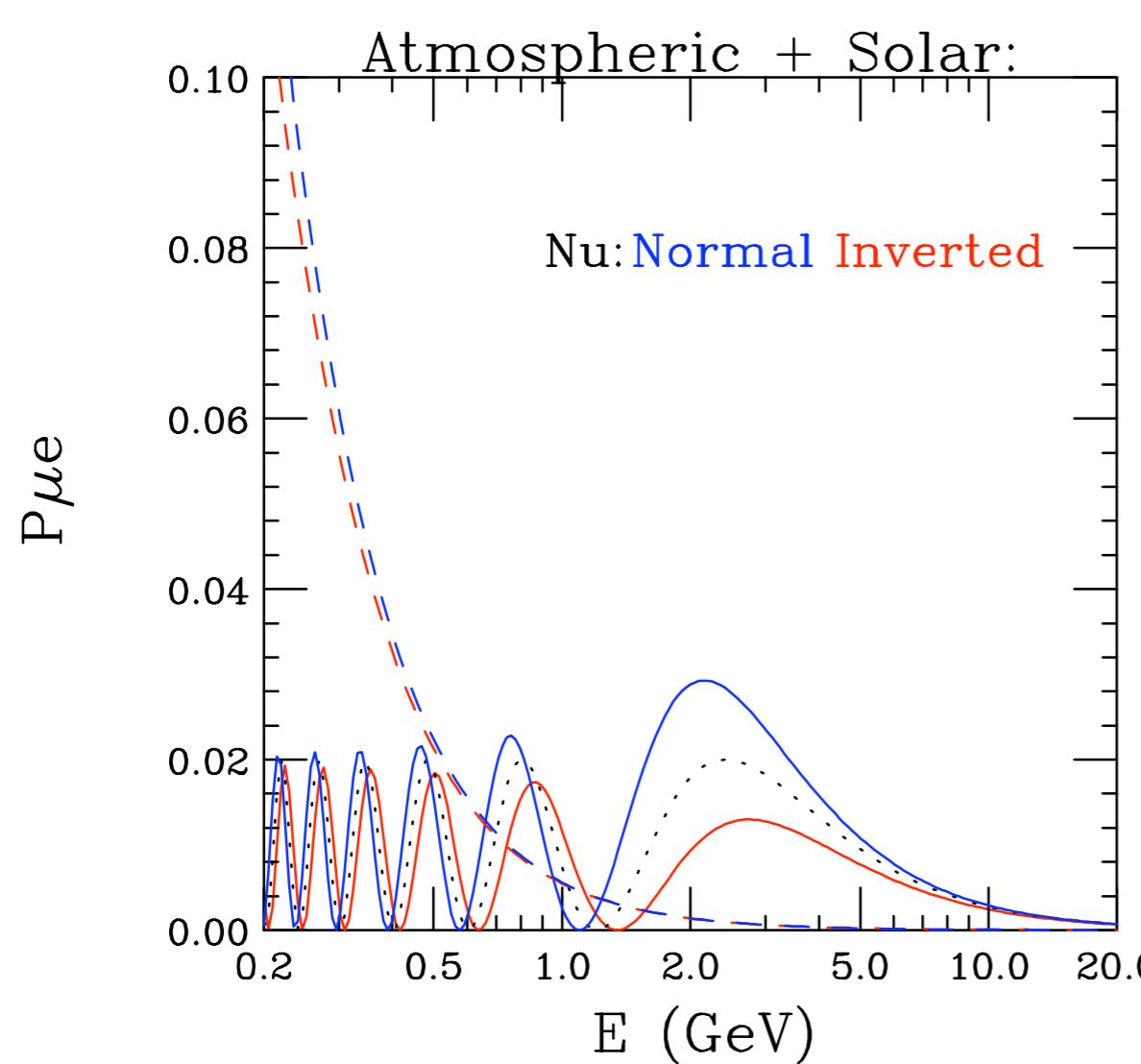
Anti-Nu: Normal Inverted

dashes $\delta = \pi/2$

solid $\delta = 3\pi/2$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

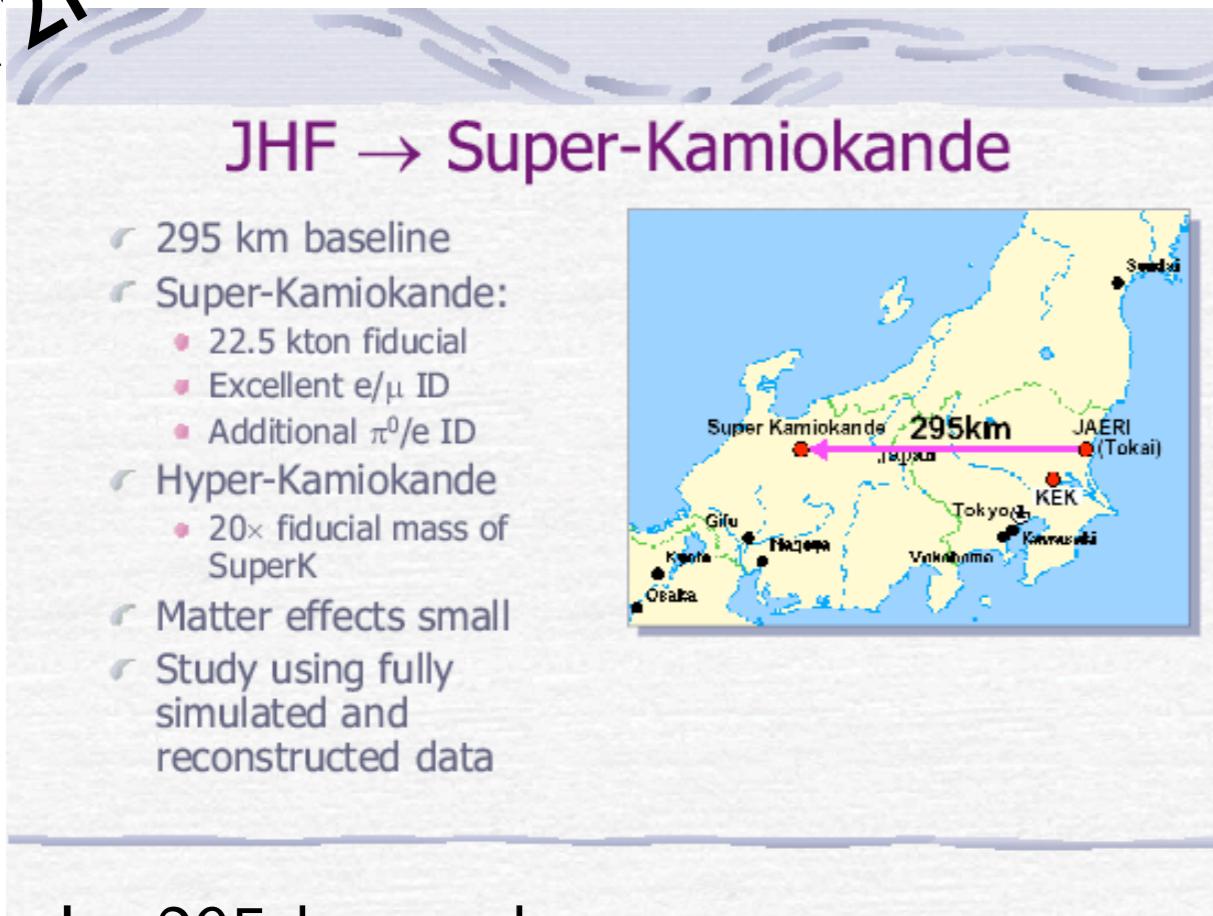


Off-Axis Beams

BNL 1994

π^0 suppression

T2K

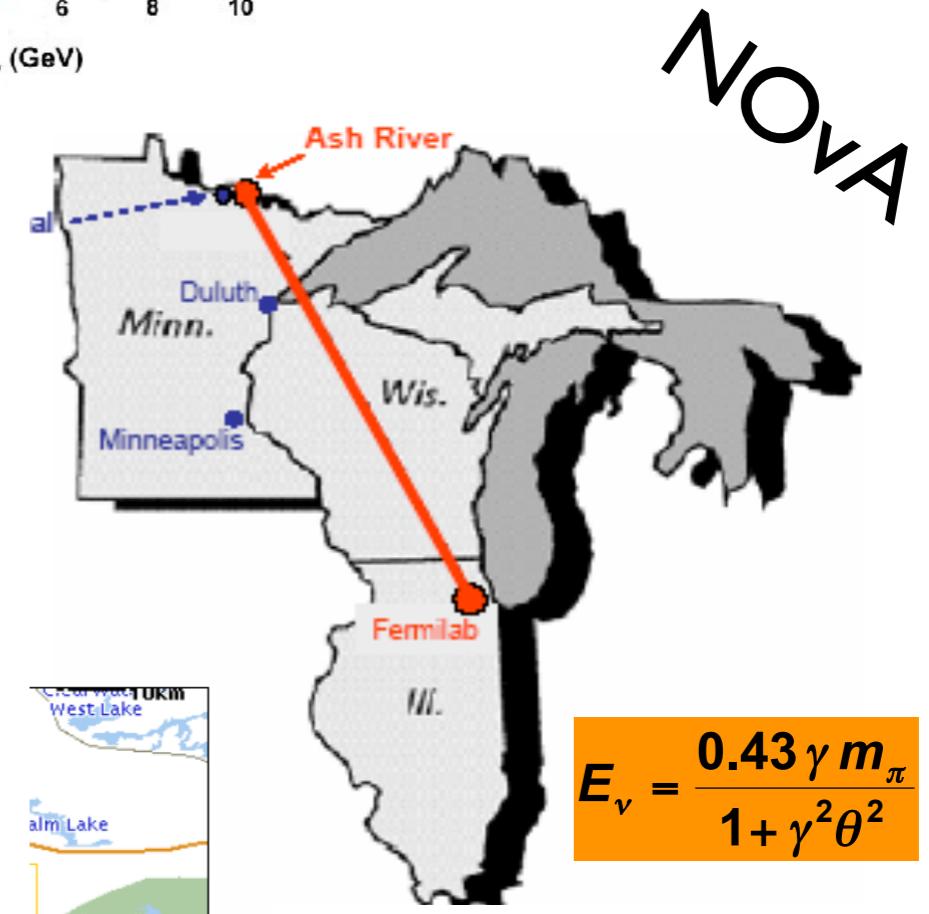
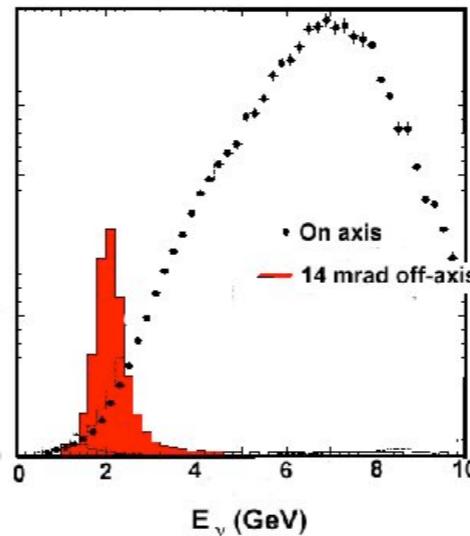


L=295 km and

Energy at Vac. Osc. Max. (vom)

$$E_{vom} = 0.6 \text{ GeV} \left\{ \frac{\delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right\}$$

0.75 upgrade to 4 MW



L=700 - 1000 km and
Energy near 2 GeV

$$E_{vom} = 1.8 \text{ GeV} \left\{ \frac{\delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right\} \times \left\{ \frac{L}{820 \text{ km}} \right\}$$

0.4 upgrade to 2 MW

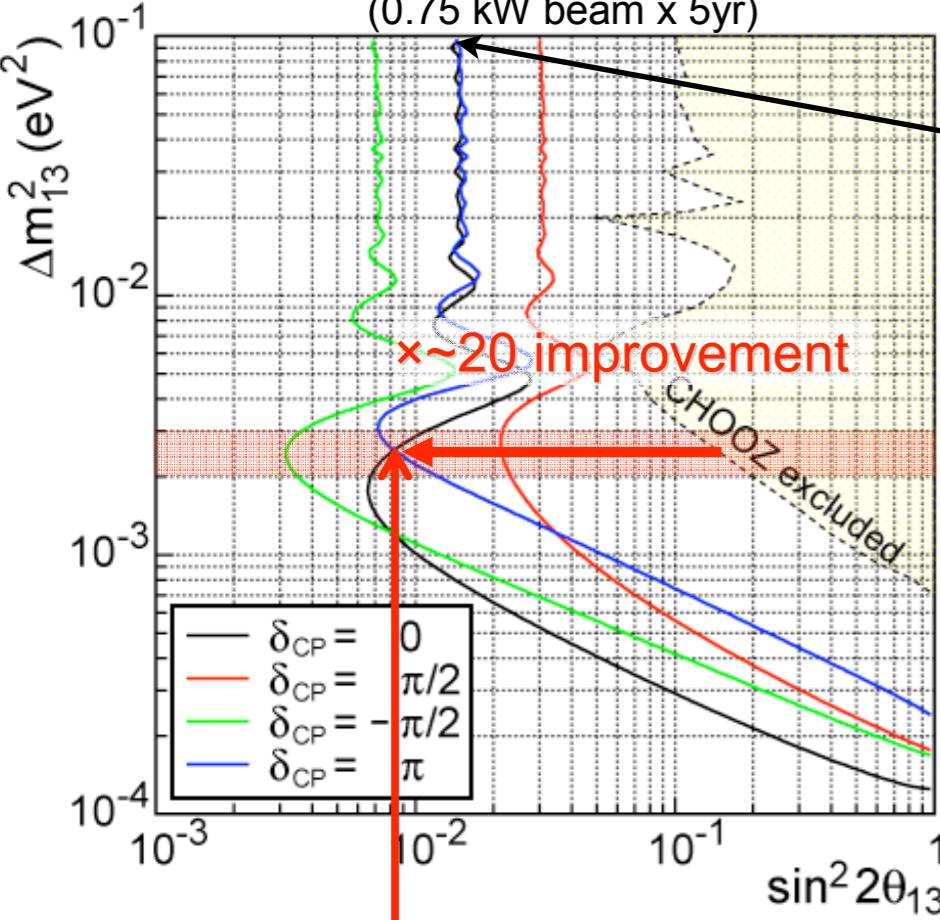
T2K

For LARGE δm_{31}^2

Search for ν_e appearance

90% C.L. sensitivity

(0.75 kW beam x 5yr)



$\sin^2 2\theta_{13} \sim 0.008$ ($\delta_{CP} = 0, \pi$)

$$\langle P(\nu_\mu \rightarrow \nu_e) \rangle = \frac{1}{2} \sin^2 \theta_{23} \sin^2 2\theta_{13} - \frac{1}{2} J \Delta_{21} + \cos^2 \theta_{23} \sin^2 2\theta_{12} \Delta_{12}^2$$

$$J = \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta$$

$$\Delta_{21} = \delta m_{21}^2 L / 4E$$

At $\delta = 0$ or π

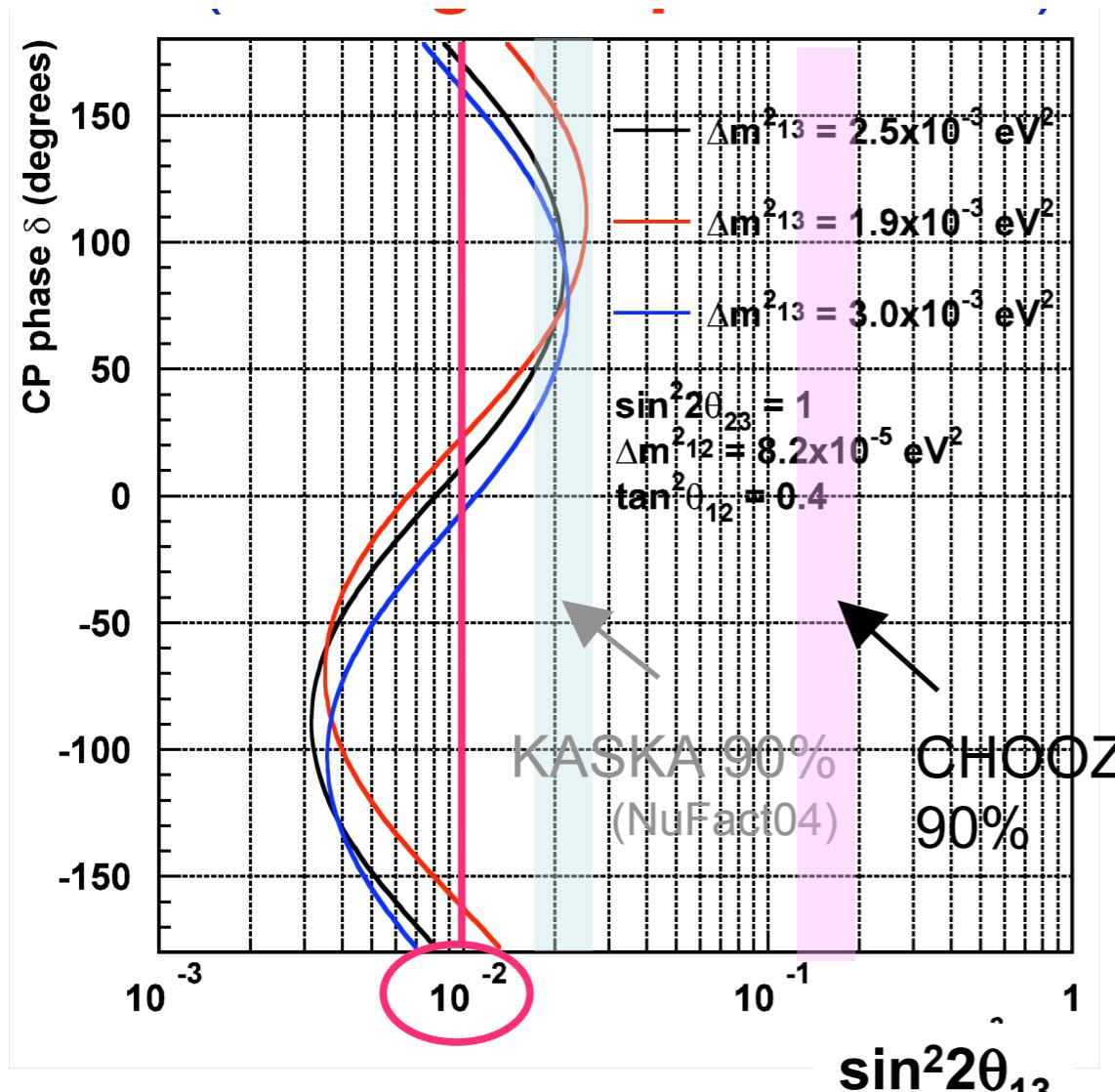
$$\begin{aligned} \langle P(\nu_\mu \rightarrow \nu_e) \rangle &= \frac{1}{2} \sin^2 \theta_{23} \sin^2 2\theta_{13} + \cos^2 \theta_{23} \sin^2 2\theta_{12} \Delta_{12}^2 \\ &\approx 0.5\% \end{aligned}$$

$$\langle P(\nu_\mu \rightarrow \nu_e) \rangle_{T2K} \approx 0.5\%$$

0.5% ν_e in beam

T2K:

Aihara for T2K, P5 talk

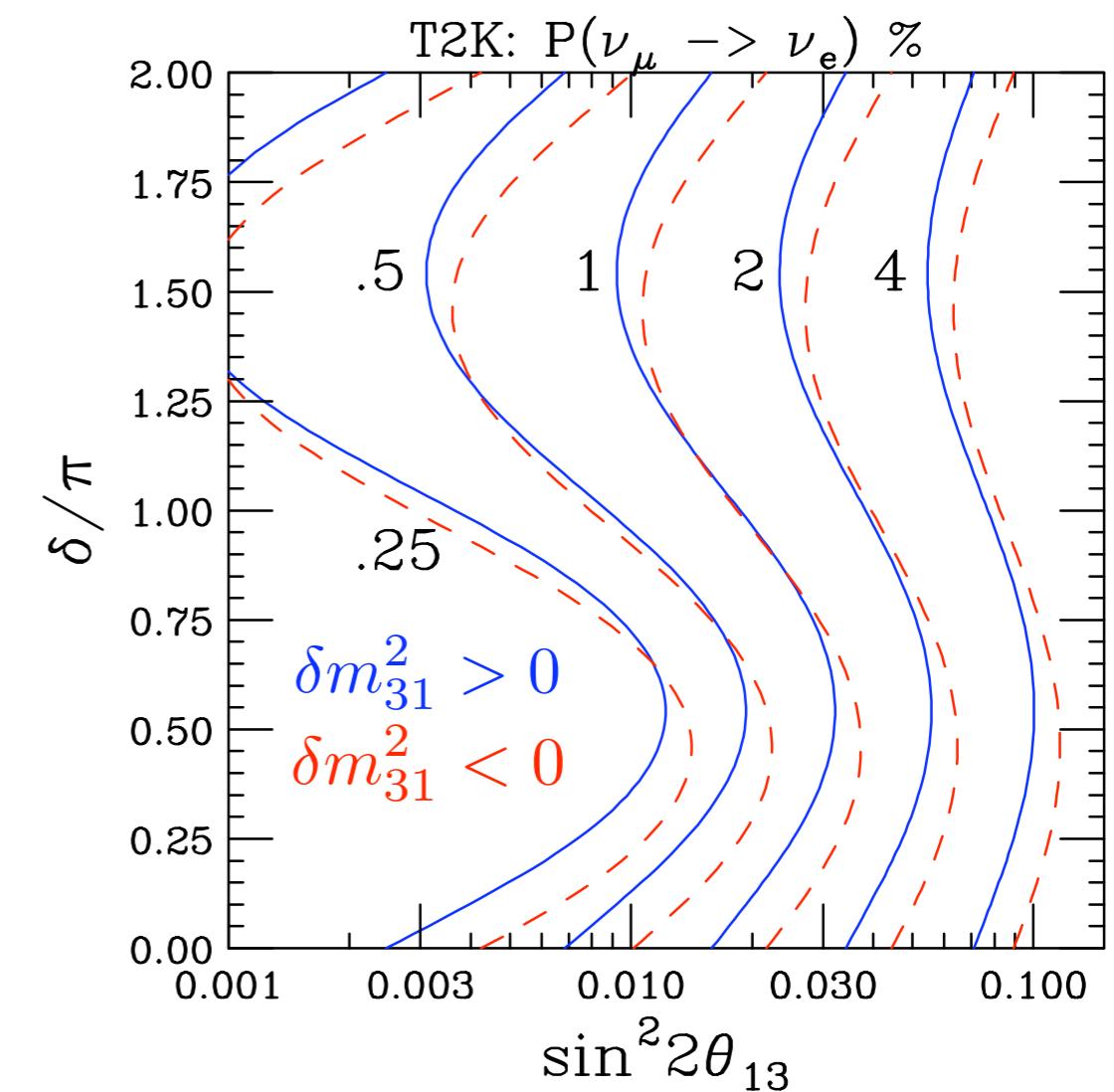
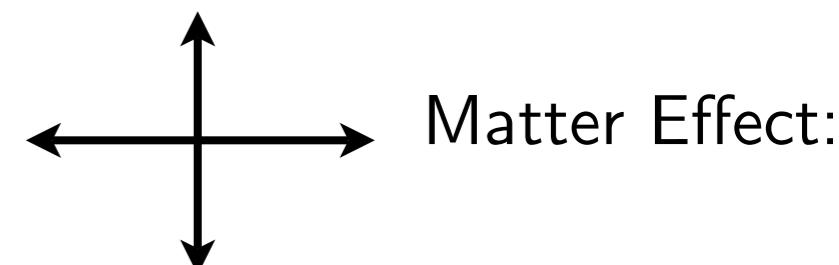


Phase I

Sensitivity approx 0.5%

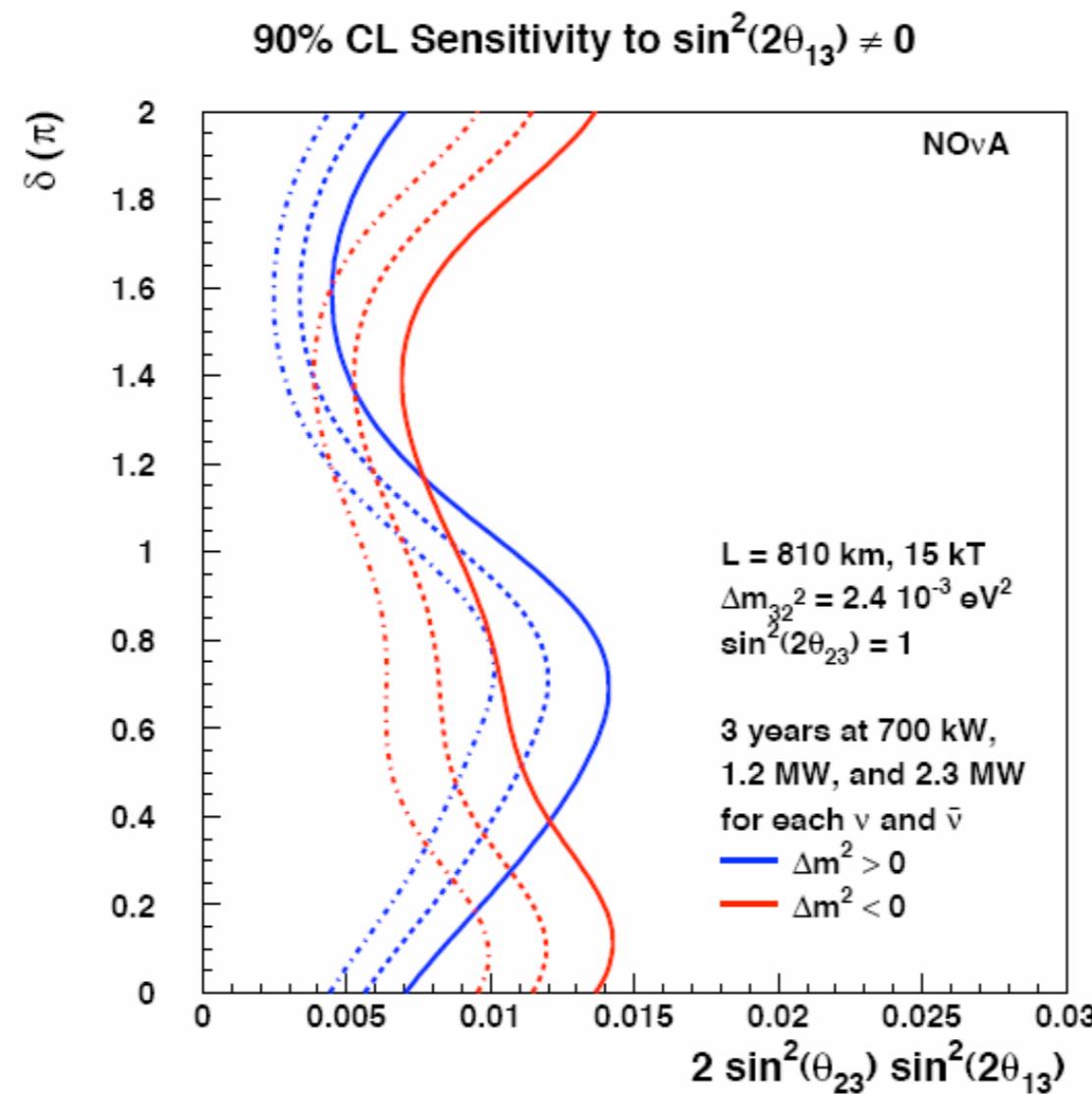
Beam 0.5%

VOM: $\Delta_{31} \neq \pi/2$



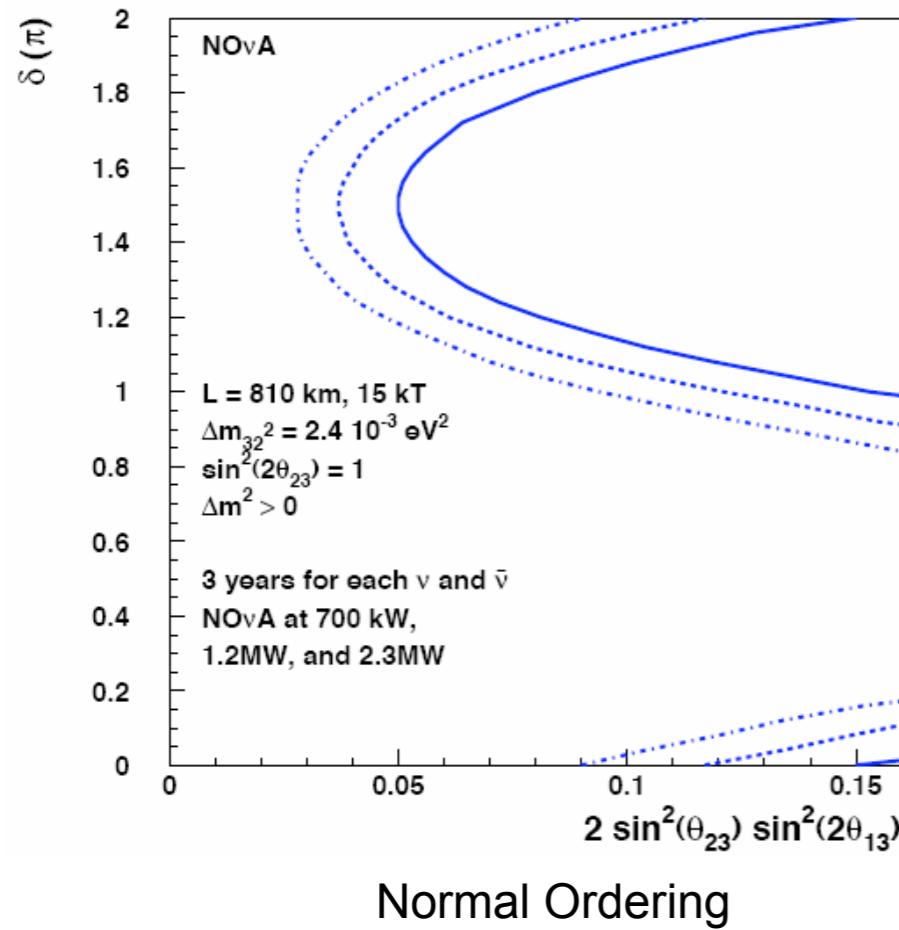


Sensitivity to $\sin^2(2\theta_{13}) \neq 0$

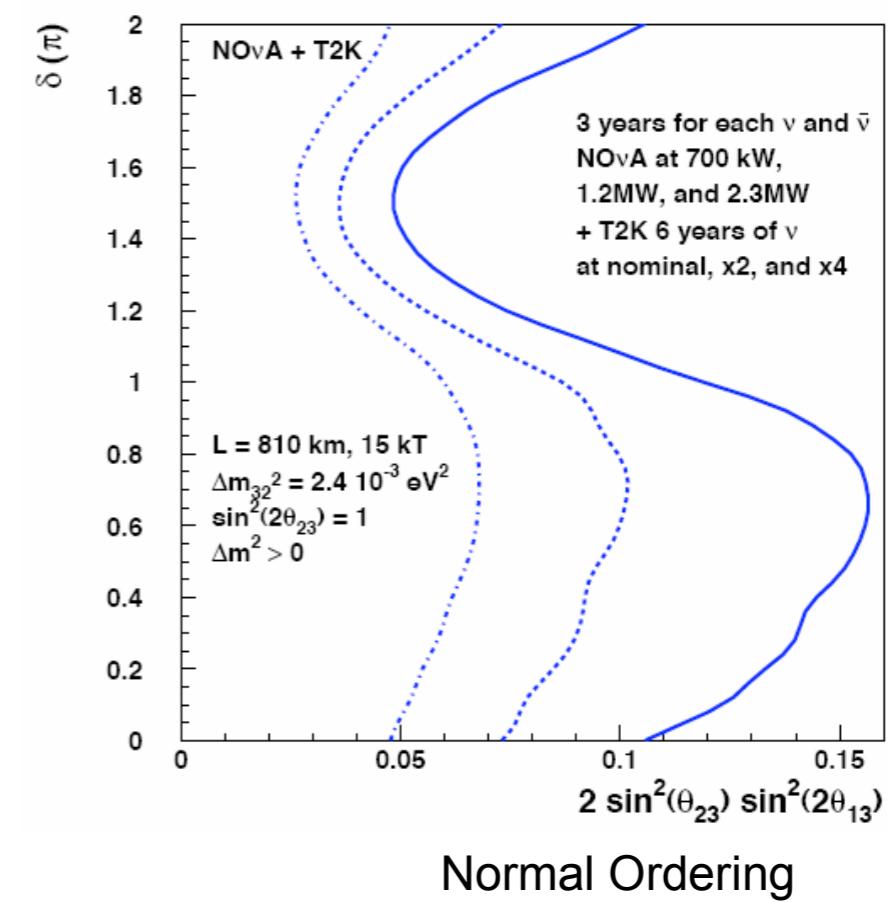




95% CL Resolution of the Mass Ordering
NO ν A Alone

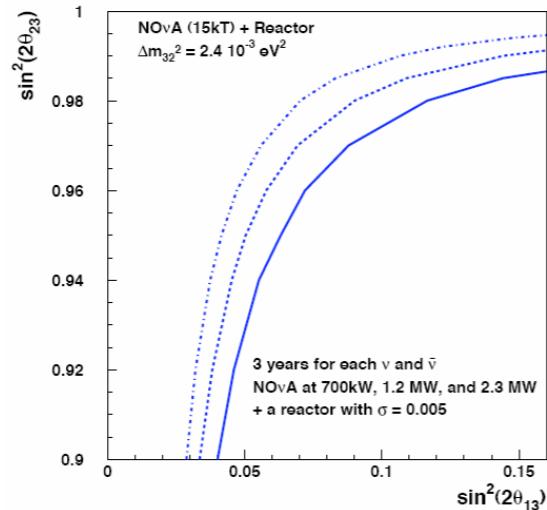


95% CL Resolution of the Mass Ordering
NO ν A Plus T2K



for Inverted Hierarchy $\delta \rightarrow \pi - \delta$

95% CL Resolution of the θ_{23} Ambiguity



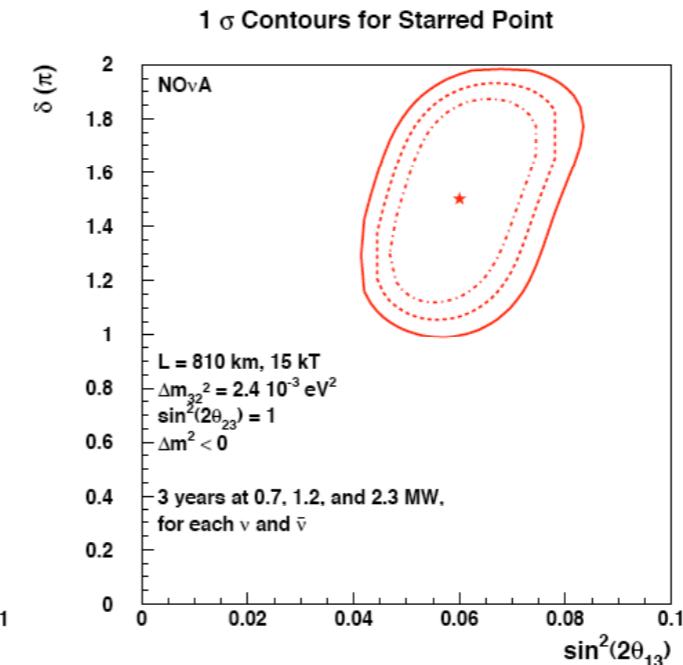
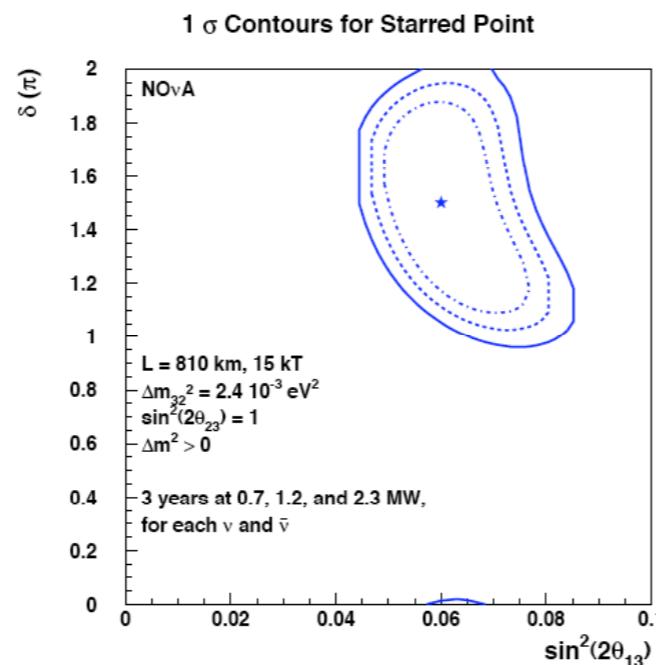
at Vac. Osc. Max.

$$P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx 2 \sin^2 \theta_{23} \sin^2 2\theta_{13} + 2P_\odot$$

in $P + \bar{P}$ the matter effects approx. cancel
and CP effects approx. cancel.



CP Violation: δ vs. θ_{13} Contours



Way Forward:

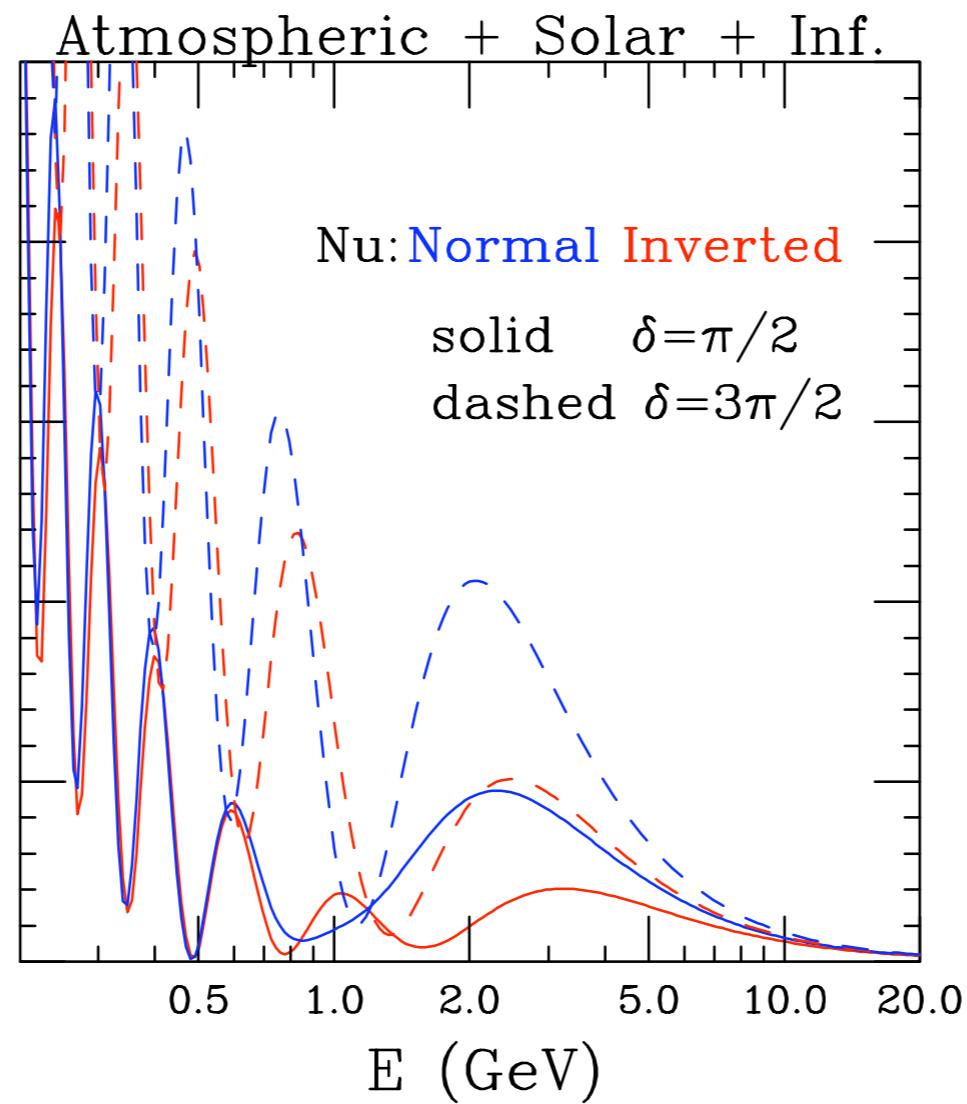
Signal Events =
Fid. Mass

- * P.O.T. (beam power*time)
- * Efficiency

2nd Oscillation Maximum

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$

Anti-Nu: Normal Inverted
dashes $\delta = \pi/2$
solid $\delta = 3\pi/2$

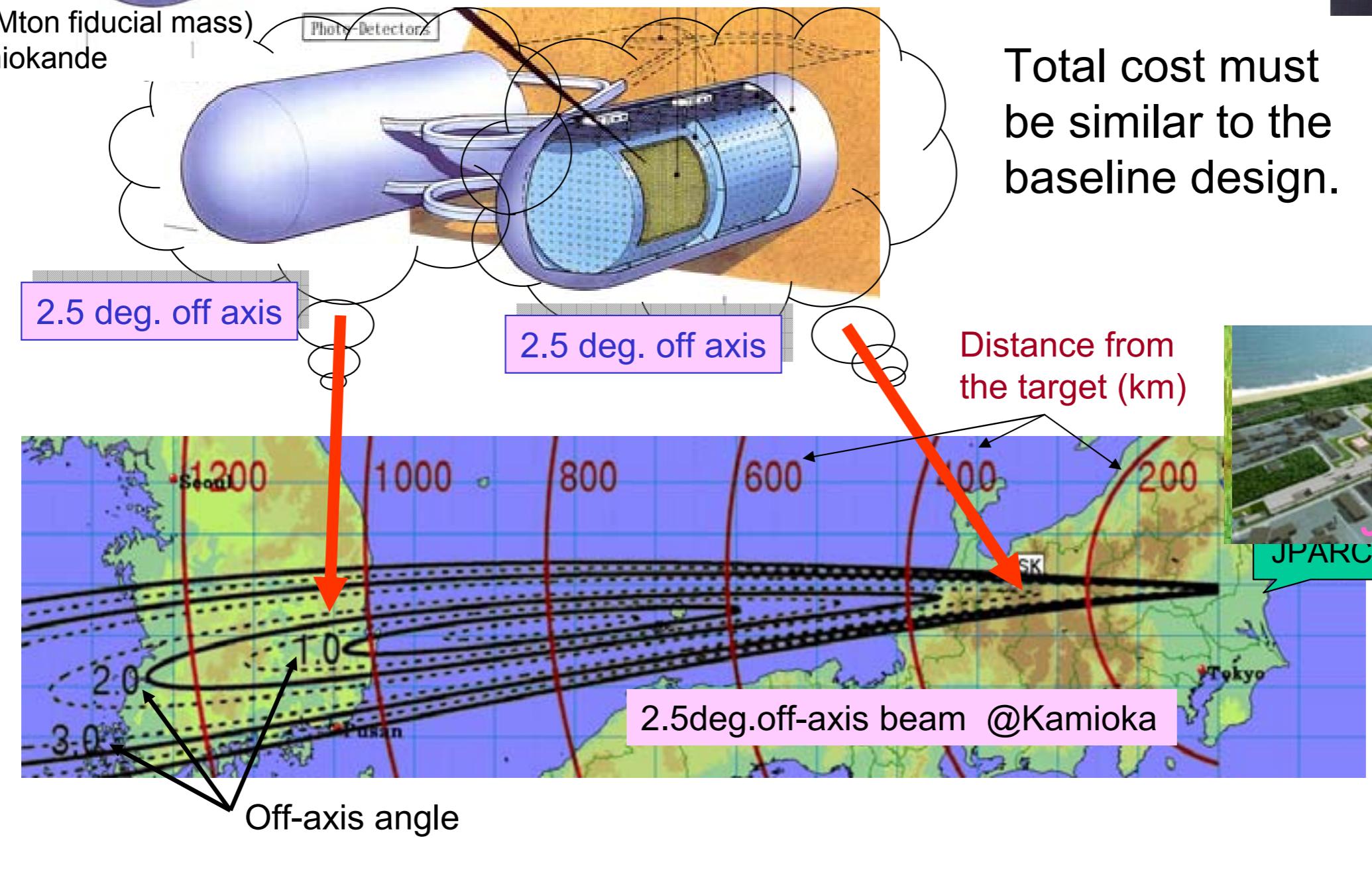


Off Axis:

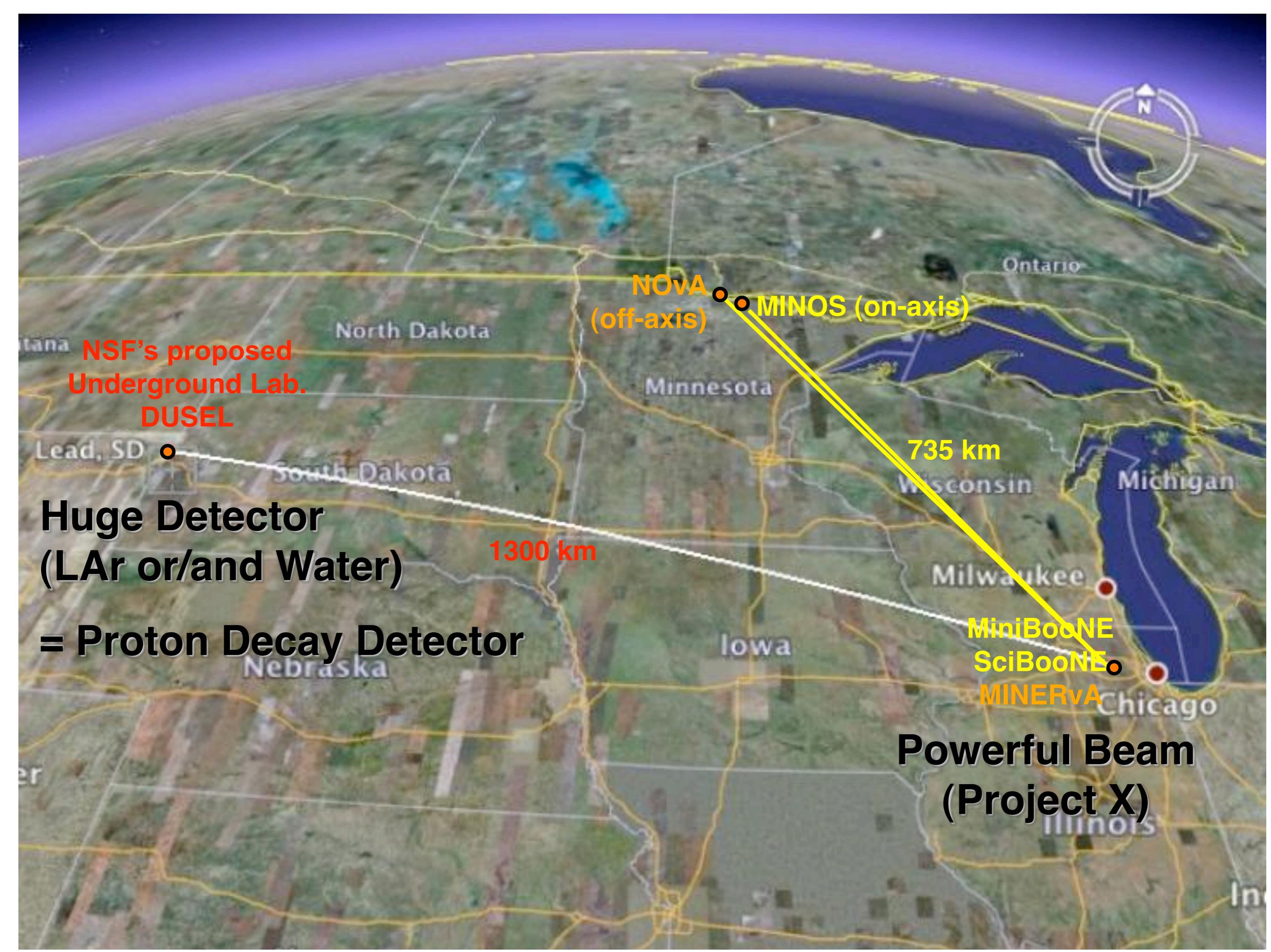


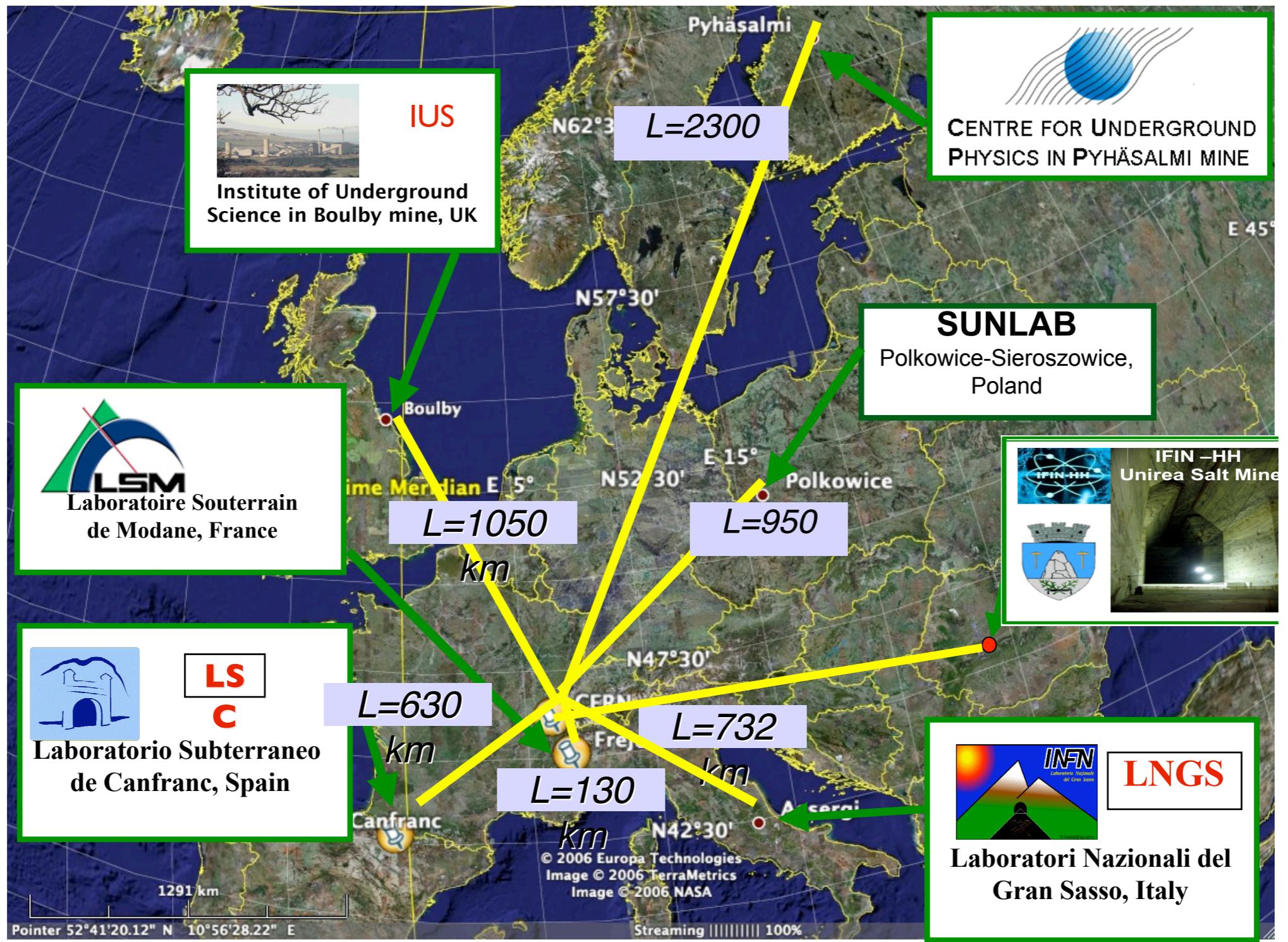
Some recent progress: detector in Korea

1Mton (0.54Mton fiducial mass)
Hyper-Kamiokande



see Kajita talk:





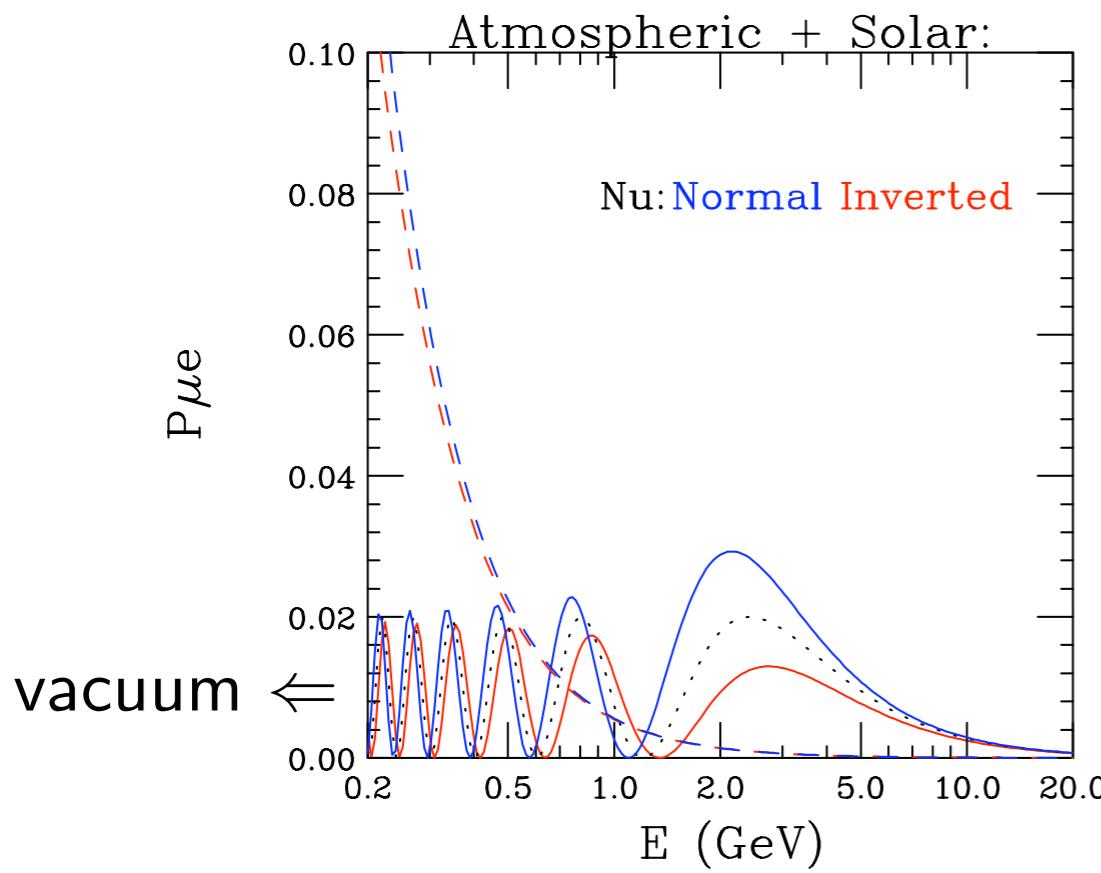
Narrow Band Beam: Same E, Longer L T2KK

Broadband Beam: Same L, Lower E Fermilab to DUSEL

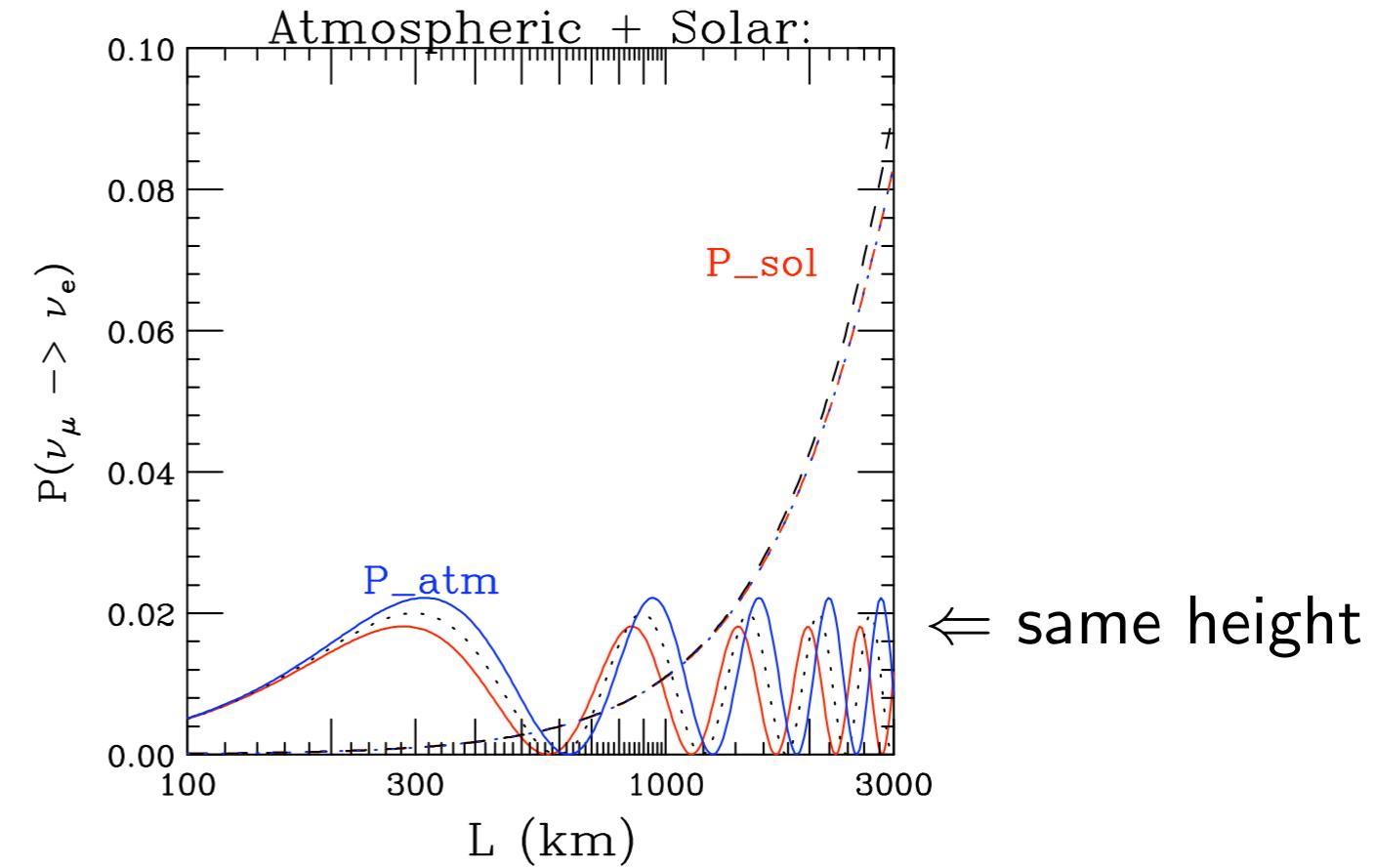
In VACUUM the SAME but NOT in MATTER

$$\sin^2 2\theta_{13} = 0.04$$

L=1200km



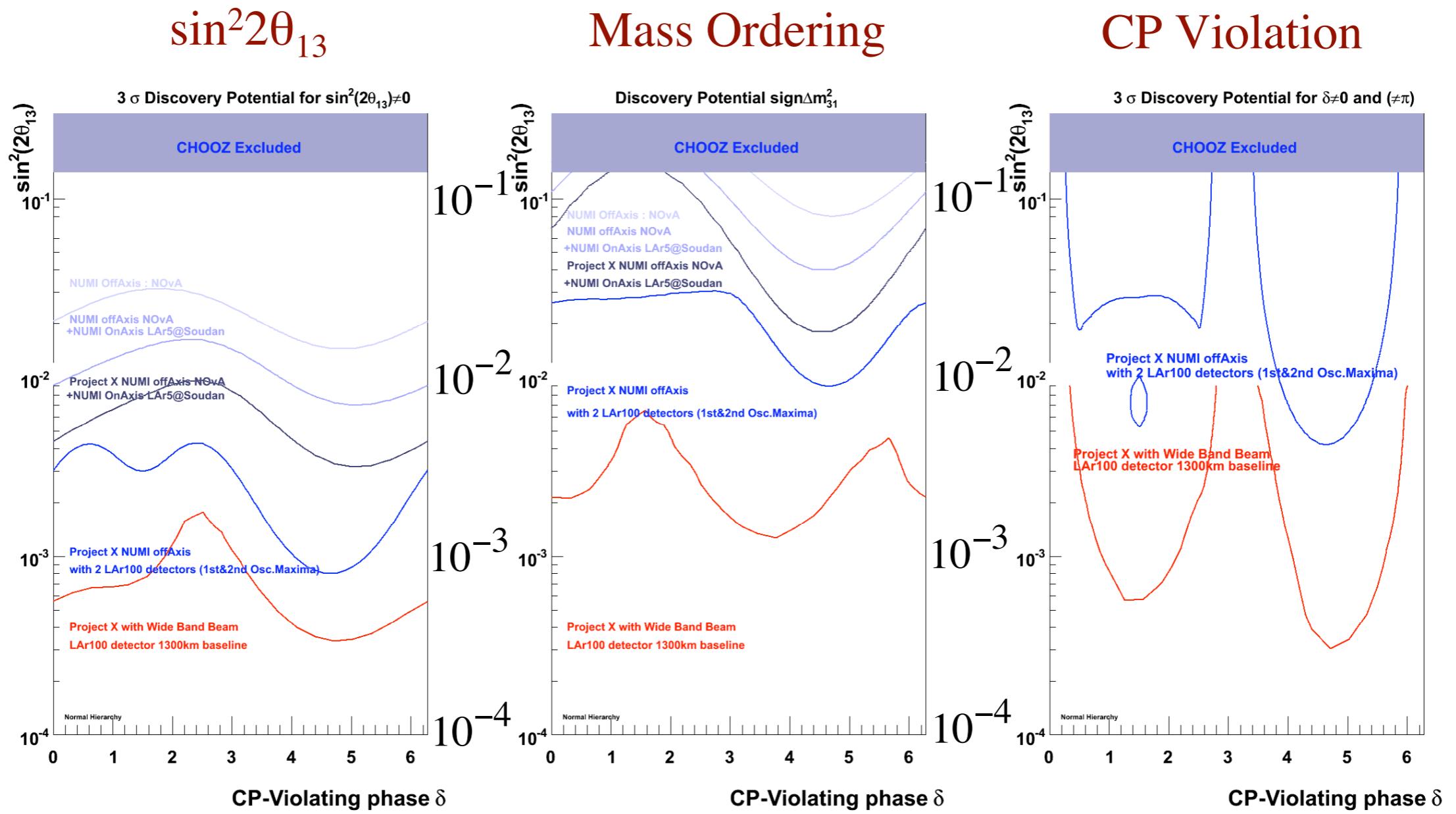
E=0.6 GeV



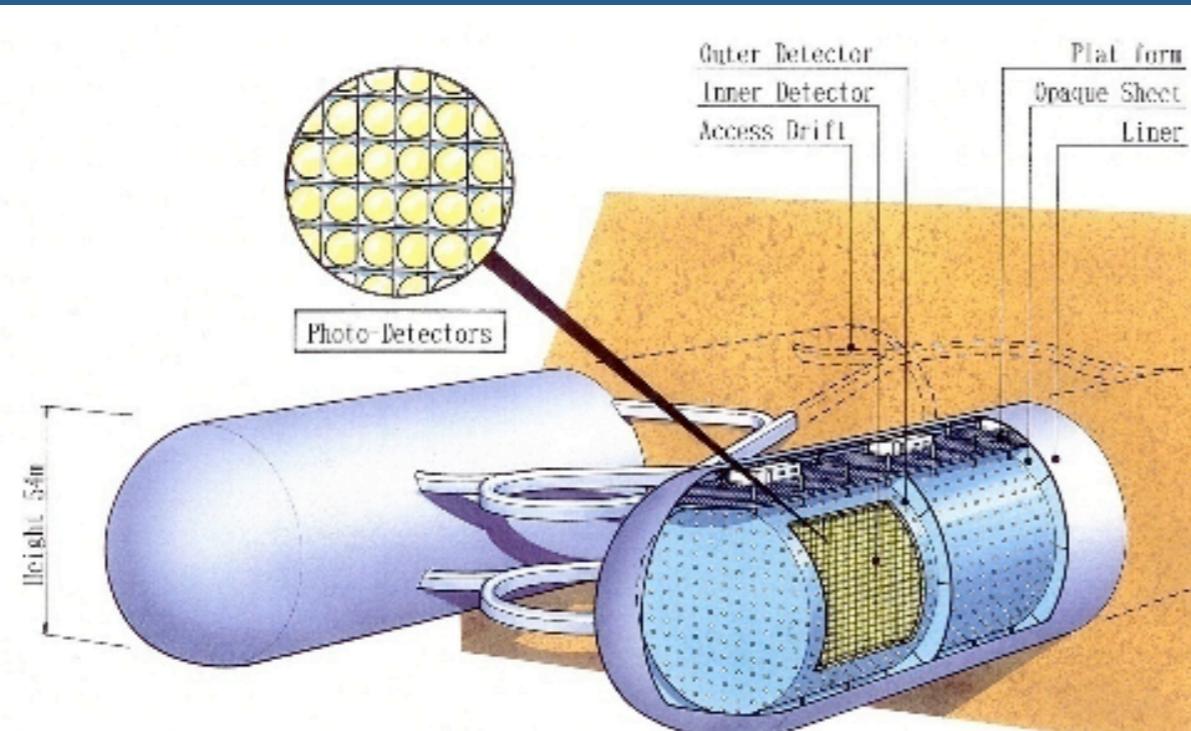
$$P_{\mu \rightarrow e} \approx | \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} |^2$$

Fermilab to DUSEL (T2KK similar)

The 3σ Reach of the Successive Phases

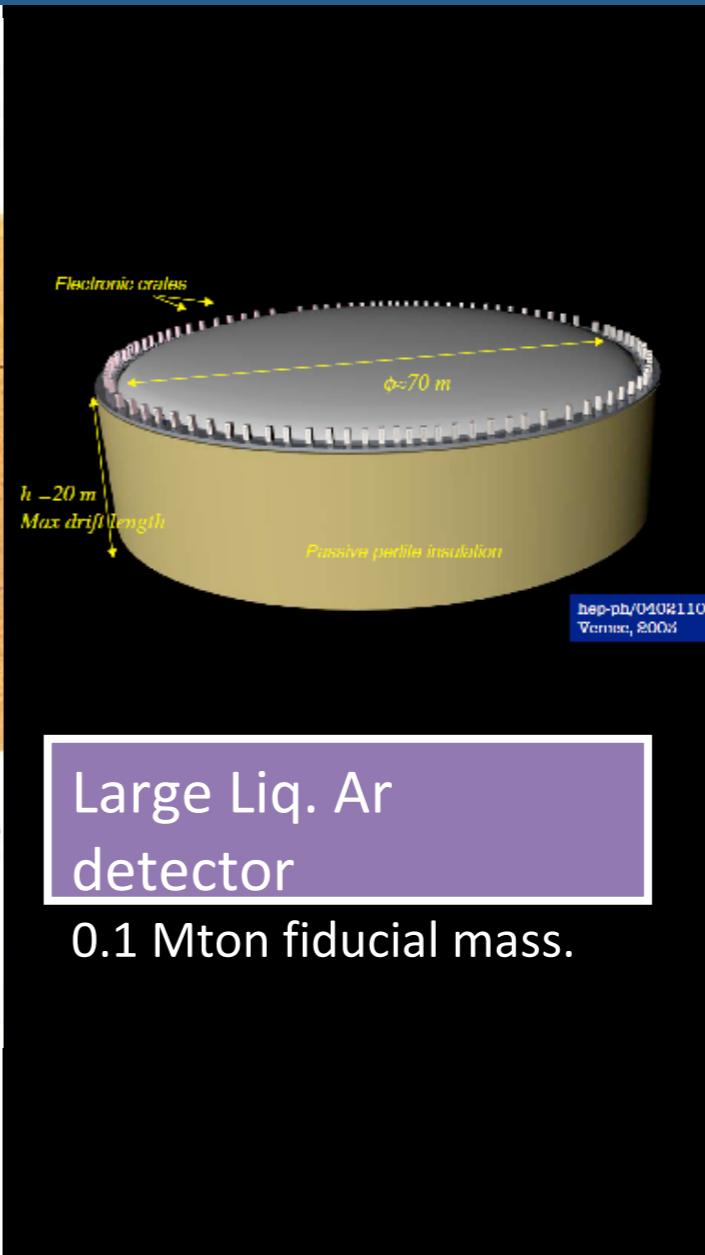


Possible detector options...



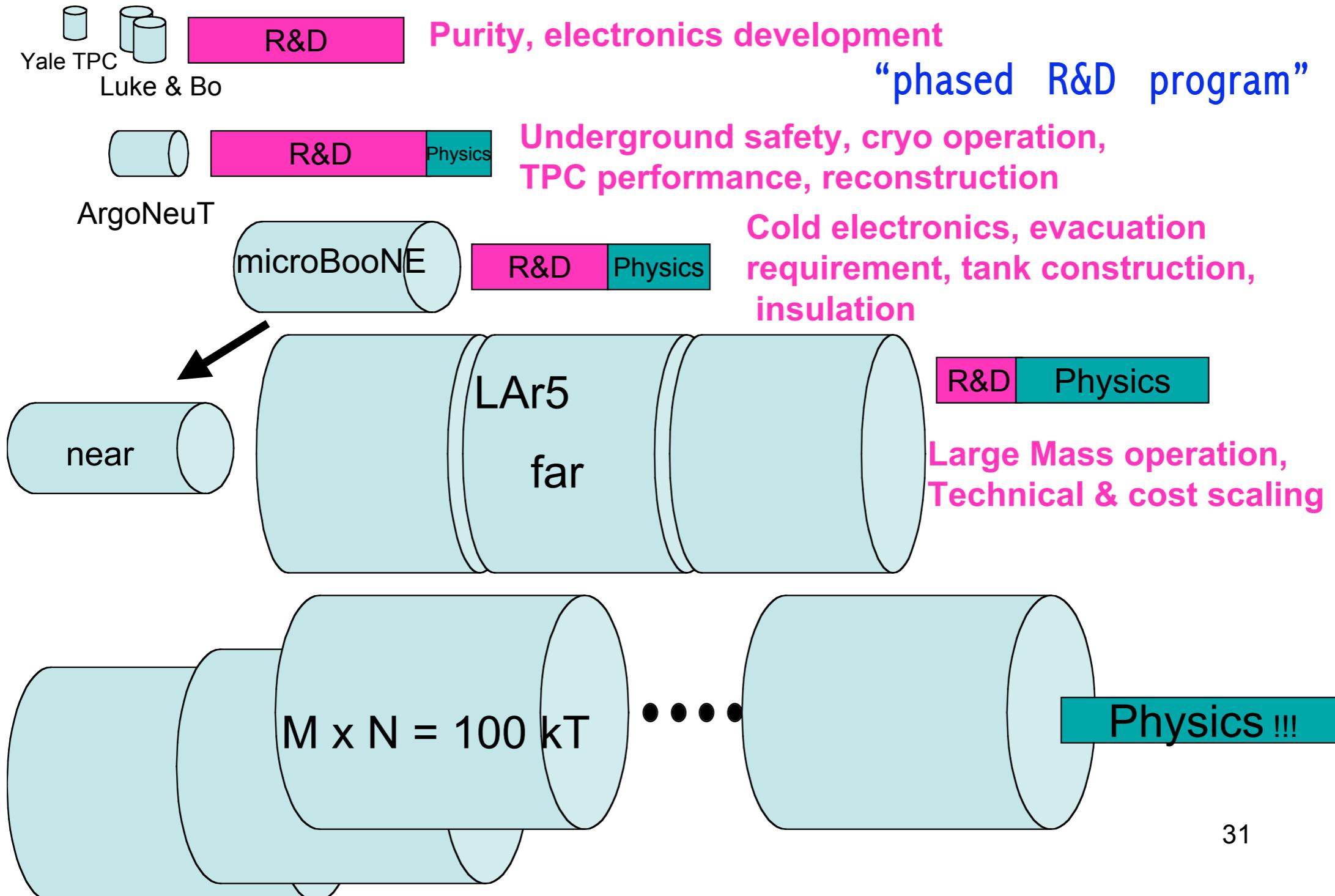
Large water Cherenkov detector
(Hyper-Kamiokande)

0.54 Mton fiducial mass (If the detector is divided into 2, located at Kamioka and Korea, fiducial mass for each unit is 0.27Mton.)



Large Liq. Ar
detector
0.1 Mton fiducial mass.

Evolution of the Liquid Argon Physics Program



Proton Decay:

$$p \rightarrow \pi^0 e^+$$

Water Cherenkov (500 kt) $\sim 10^{35}$ y (5000 kt y exposure)

Limit SK-I and II: $t > 8.4 \times 10^{33}$ y

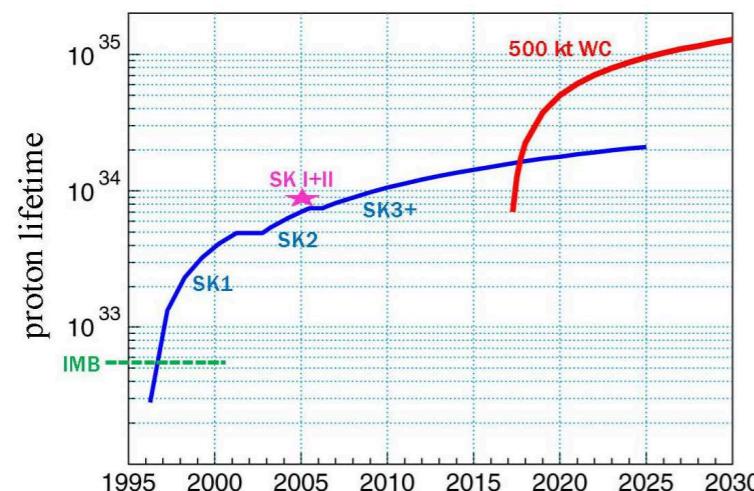
$$p \rightarrow K^+ \bar{\nu}$$

Liquid Argon (100kt) $\sim 10^{35}$ y (1000 kt y exposure)

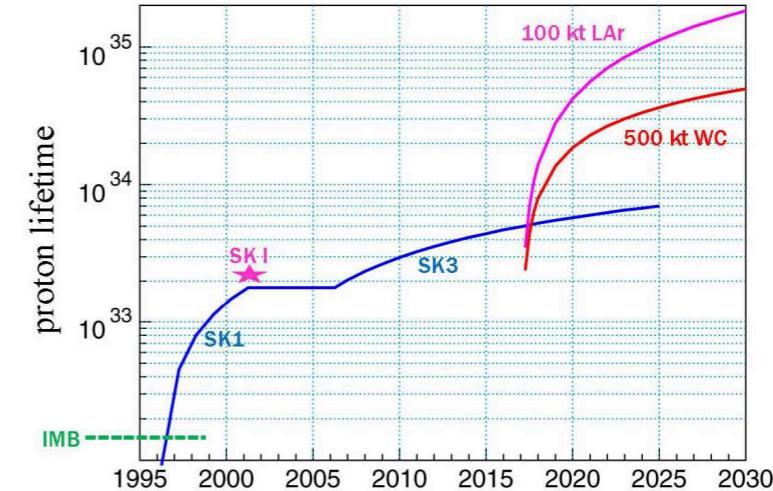
Liquid Scintillator (50kt) 5×10^{34} y (500 kt y exposure)

Limit SK-I: $t > 2.3 \times 10^{32}$ y

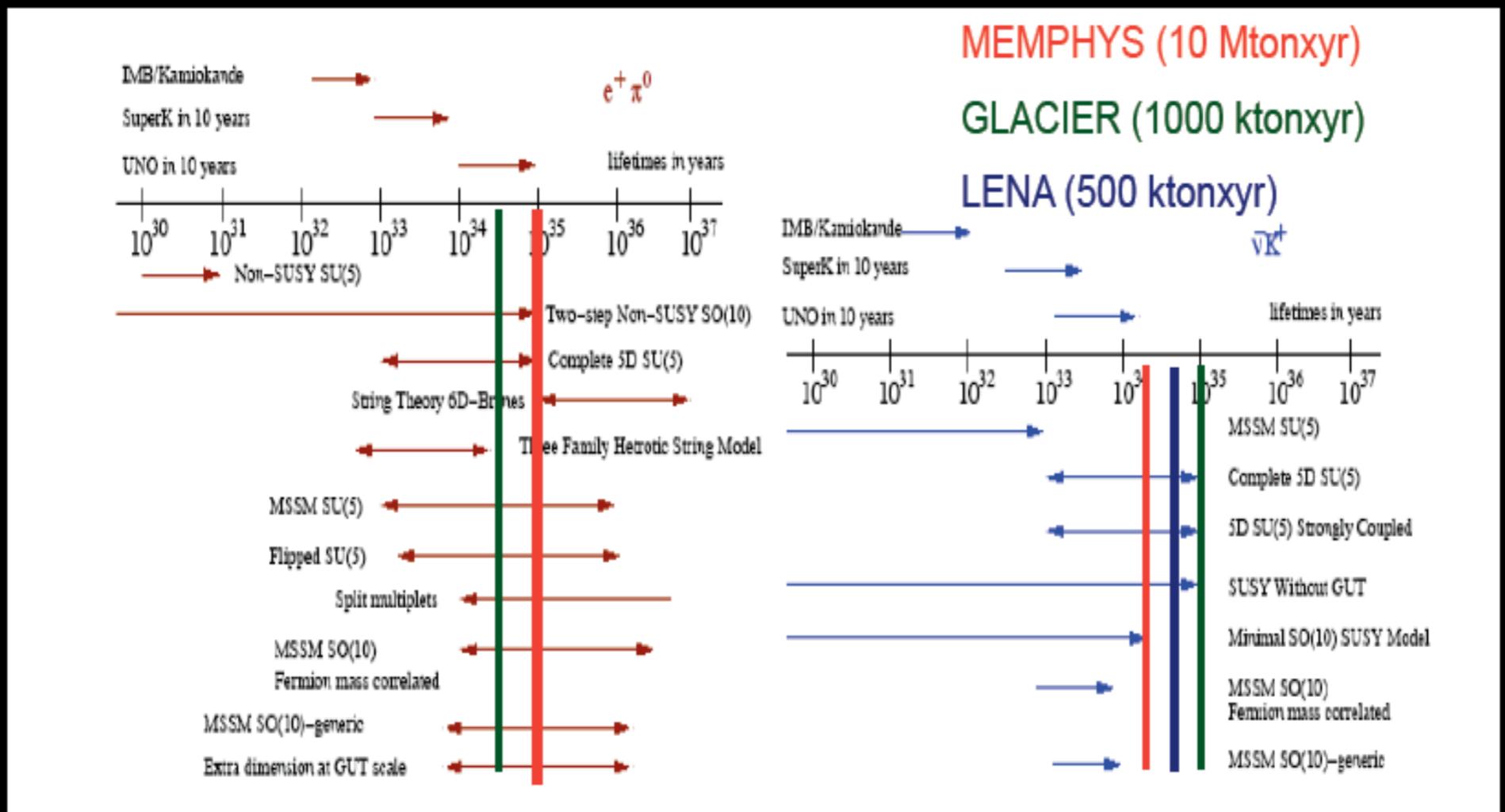
$$p \rightarrow e^+ \pi^0$$



$$p \rightarrow K^+ \nu$$



Sensitivity to proton decay: comparison with theory



Higher dimension models (eg. 6D SO(10)) not included

LAGUNA

SUPERNOMA:



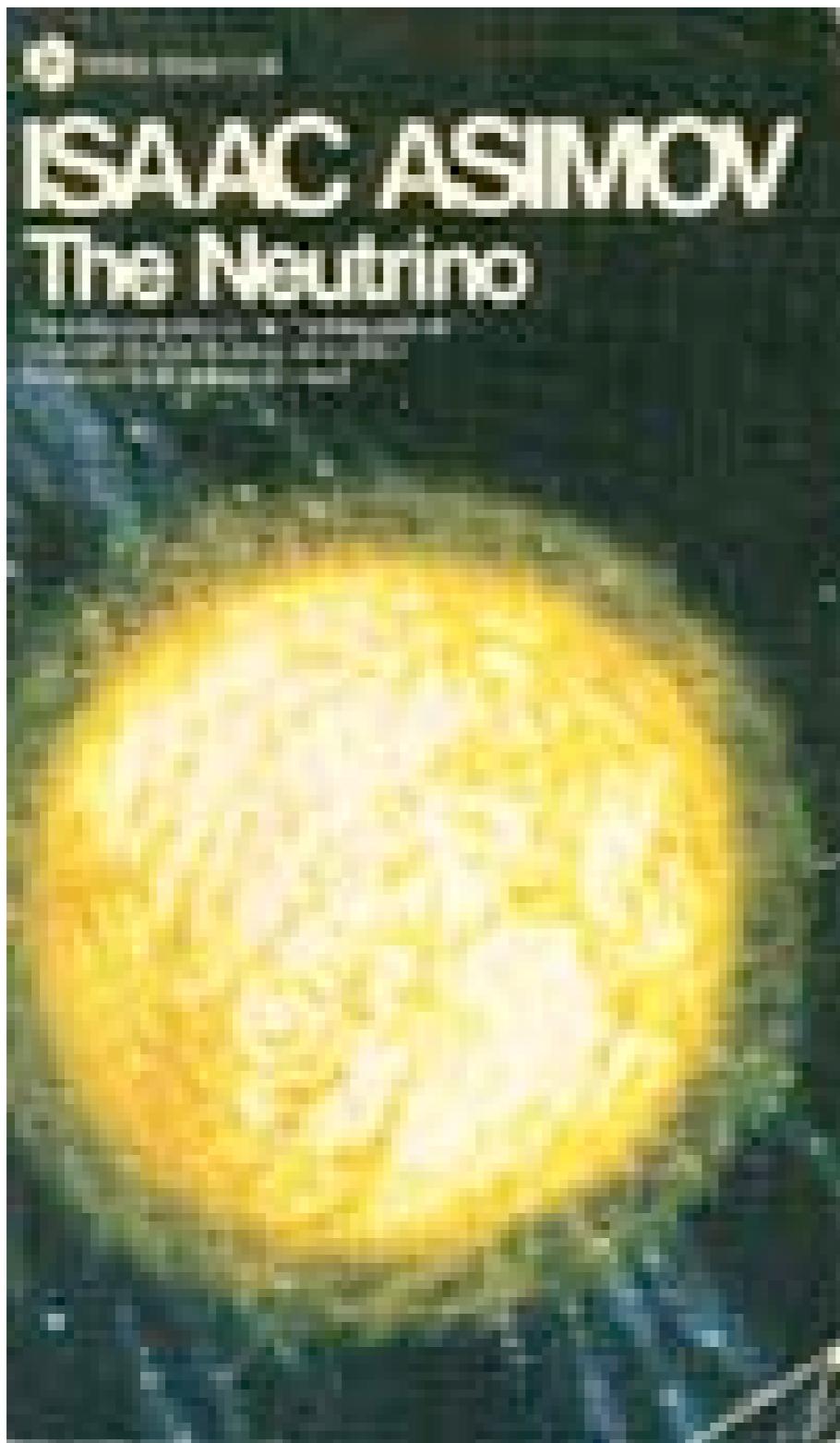
© Anglo-Australian Observatory



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1987a Feb 24

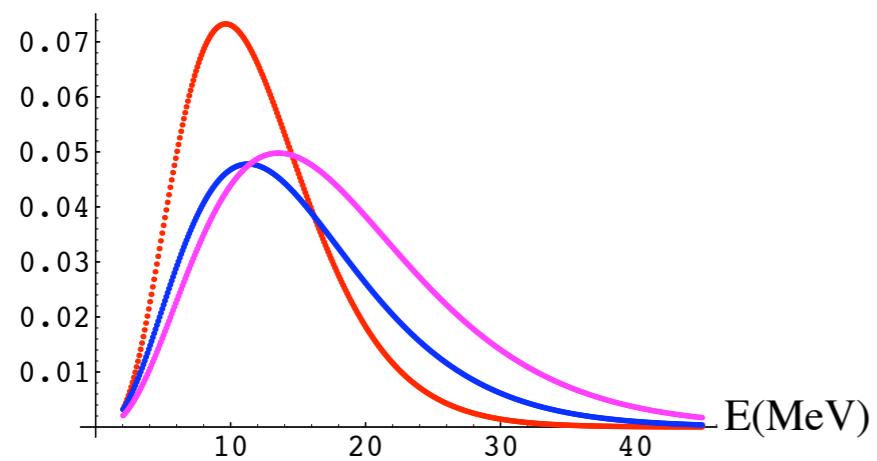
1966



And yet the nothing-particle is not a nothing at all

We are “due” for a supernova anytime now we can only hope that it will hold off until the science of neutrino astronomy is further advanced.

Neutrino fluxes:

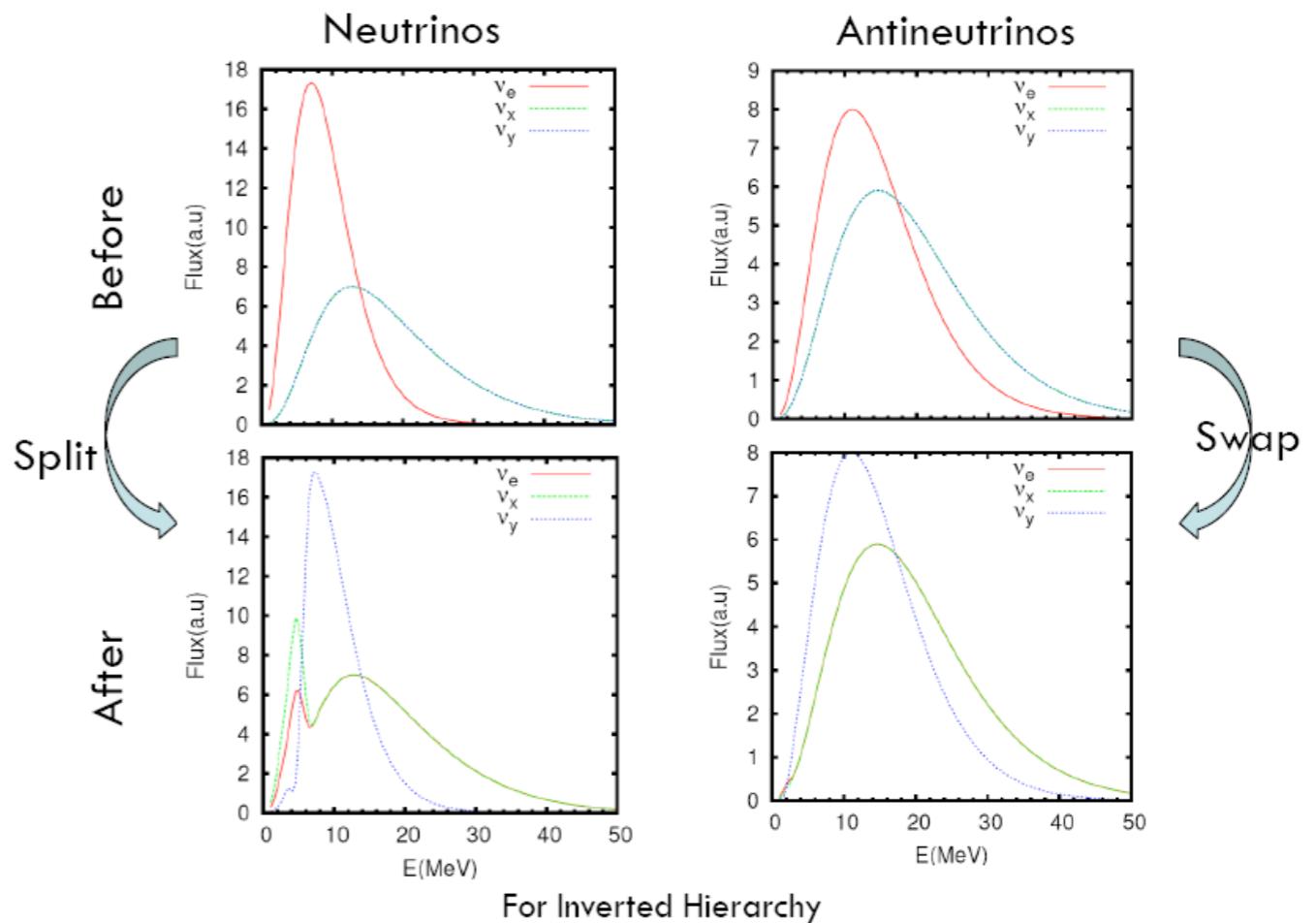


$$E_0(\nu_e) \approx 10\text{--}12 \text{ MeV}$$

$$E_0(\bar{\nu}_e) \approx 13\text{--}16 \text{ MeV}$$

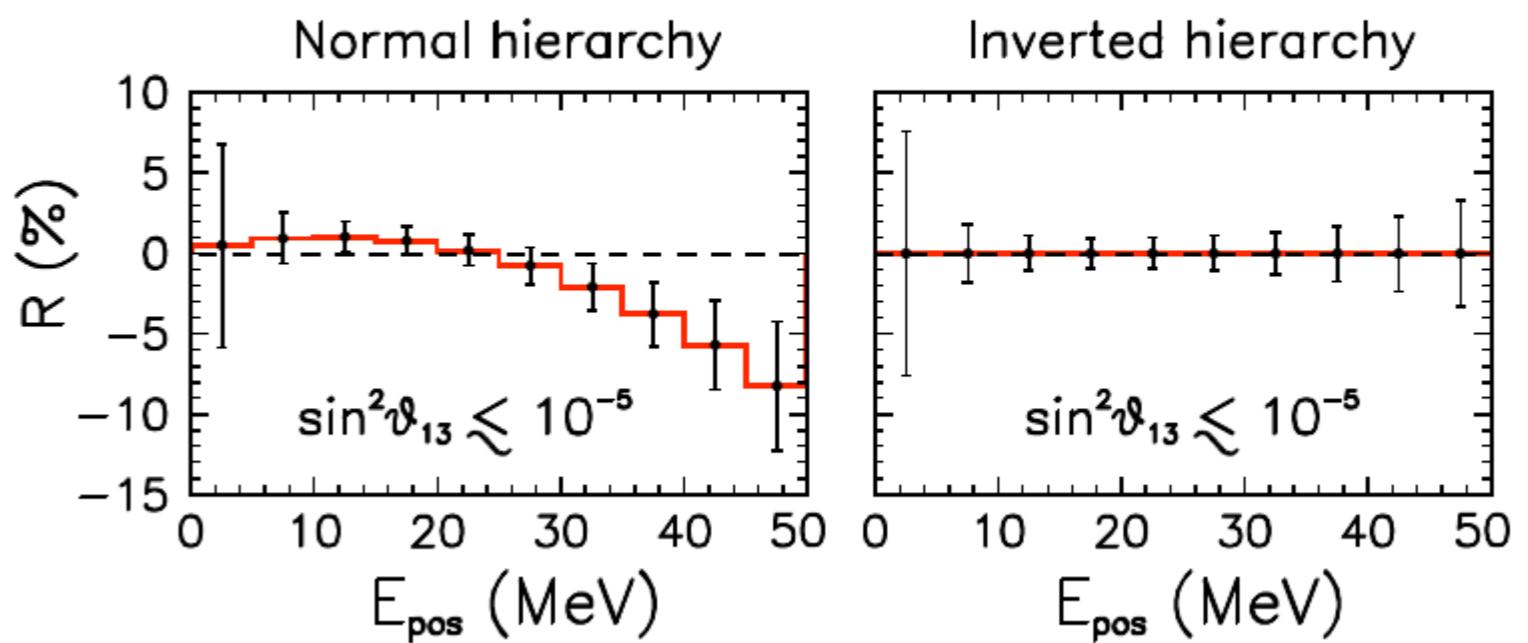
$$E_0(\nu_x) \approx 15\text{--}25 \text{ MeV}$$

SN spectra reaching Earth



Mass hierarchy at very small θ_{13}

- $R = (N_{\text{Shadowed}} - N_{\text{Not Shadowed}}) / N_{\text{Not Shadowed}}$.



Conclusions

- A Large Underground Detector is the way forward for:
 - Neutrino Properties
($\sin^2 \theta_{13}$, sign{ δm_{31}^2 }, $\sin \delta$, $\sin^2 \theta_{23}$, surprises)
 - Proton Decay
(beyond 10^{35} years in both $p \rightarrow \pi^0 e^+$ and $p \rightarrow K^+ \bar{\nu}$)
 - Supernova Neutrinos
($> 10^5$ events for galactic event)
 - Other - solar, atmospheric neutrinos ...
- Both Japan (T2HK, T2KK ...) and Europe (LAGUNA) are exploring similar options.